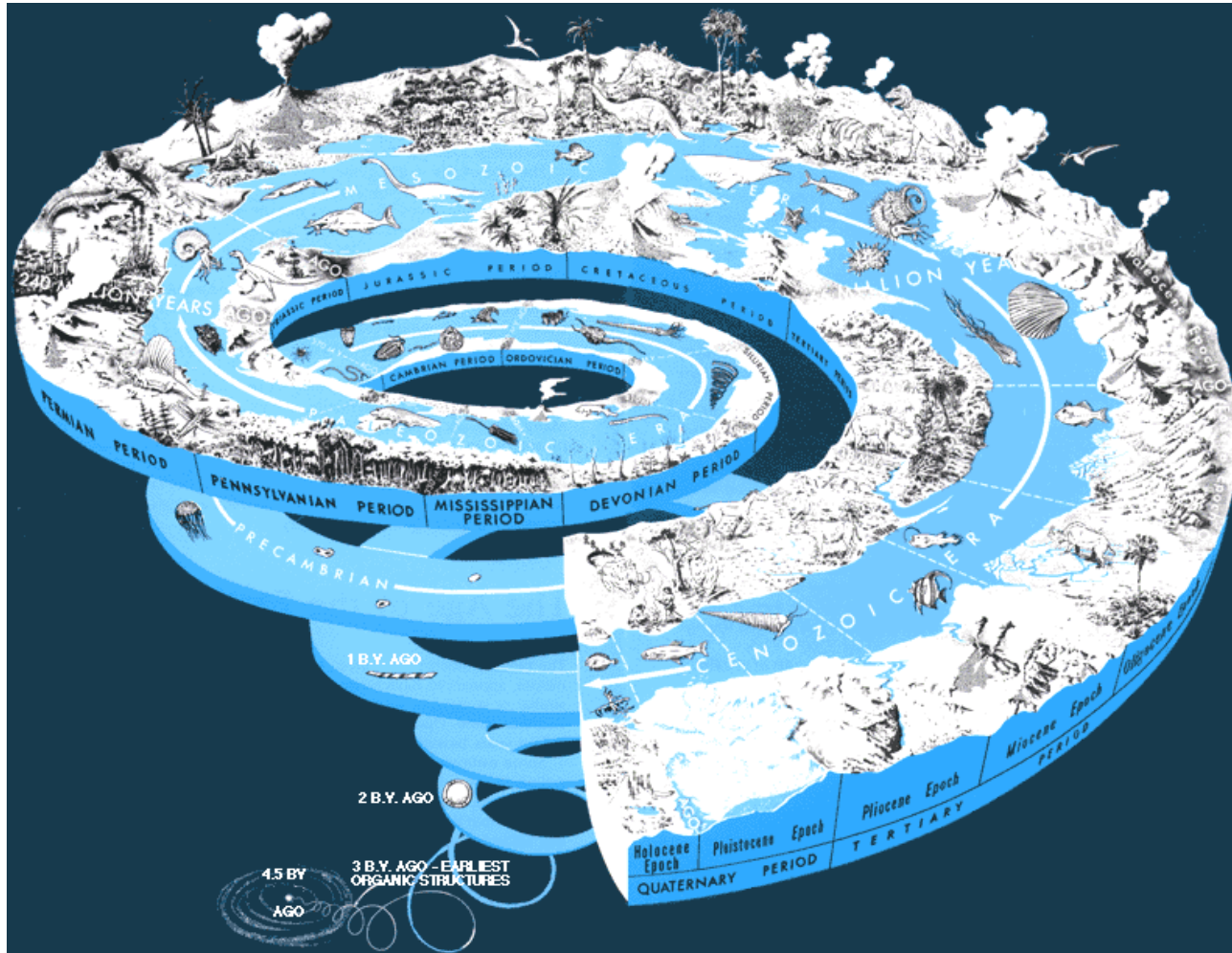


# Isotopengeochemie Geochronologie



# Radiogene Isotopensysteme

Goldschmidt-Klassifikation im Periodensystem der Elemente

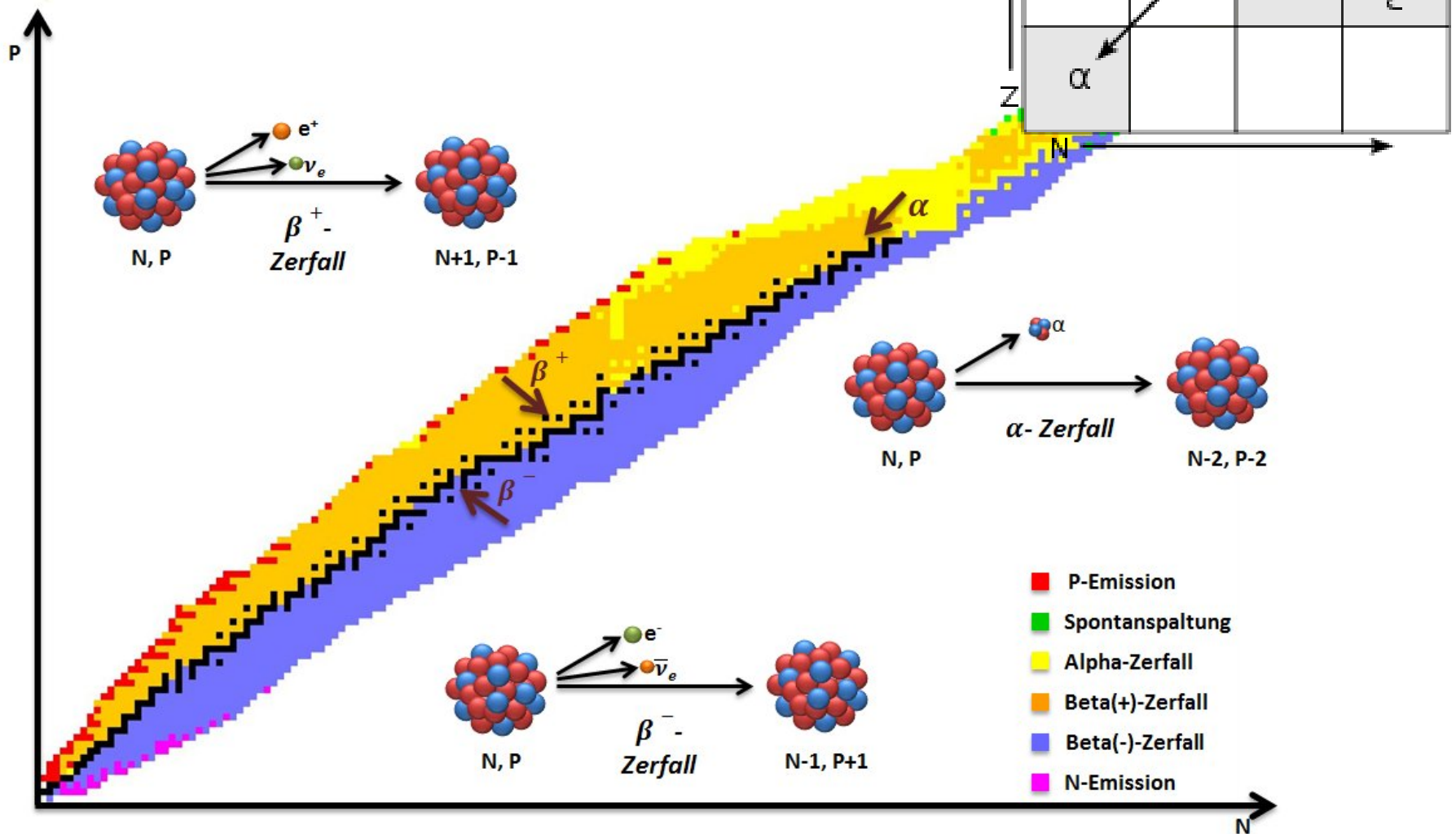
|             |    |    |     |    |      |      |       |      |      |      |      |       |       |       |       |     |     |     |     |
|-------------|----|----|-----|----|------|------|-------|------|------|------|------|-------|-------|-------|-------|-----|-----|-----|-----|
| 1           | 1  | 2  |     |    |      |      |       |      |      |      |      |       | 13    | 14    | 15    | 16  | 17  | 18  |     |
| 1           | H  |    |     |    |      |      |       |      |      |      |      |       |       |       |       |     |     | He  |     |
| 2           | Li | Be |     |    |      |      |       |      |      |      |      |       | B     | C     | N     | O   | F   | Ne  |     |
| 3           | Na | Mg | 3   | 4  | 5    | 6    | 7     | 8    | 9    | 10   | 11   | 12    | Al    | Si    | P     | S   | Cl  | Ar  |     |
| 4           | K  | Ca | Sc  | Ti | V    | Cr   | Mn    | Fe   | Co   | Ni   | Cu   | Zn    | Ga    | Ge    | As    | Se  | Br  | Kr  |     |
| 5           | Rb | Sr | Y   | Zr | Nb   | Mo   | (43)  | Tc   | Ru   | Rh   | Pd   | Ag    | Cd    | In    | Sn    | Sb  | Te  | I   | Xe  |
| 6           | Cs | Ba | Lan | Hf | Ta   | W    | 75    | Re   | Os   | Ir   | Pt   | Au    | Hg    | Tl    | Pb    | Bi  | Po  | At  | Rn  |
| 7           | Fr | Ra | Act | Rf | Db   | Sg   | (107) | Bh   | Hs   | Mt   | Ds   | Rg    | Cn    | Uut   | Uuq   | Uup | Uuh | Uus | Uuo |
| Lanthanoide | 57 | 58 | 59  | 60 | (61) | 62   | 63    | 64   | 65   | 66   | 67   | 68    | 69    | 70    | 71    |     |     |     |     |
|             | La | Ce | Pr  | Nd | Pm   | Sm   | Eu    | Gd   | Tb   | Dy   | Ho   | Er    | Tm    | Yb    | Lu    |     |     |     |     |
| Actinoide   | 89 | 90 | 91  | 92 | (93) | (94) | (95)  | (96) | (97) | (98) | (99) | (100) | (101) | (102) | (103) |     |     |     |     |
|             | Ac | Th | Pa  | U  | Np   | Pu   | Am    | Cm   | Bk   | Cf   | Es   | Fm    | Md    | No    | Lr    |     |     |     |     |

Legende:

|              |            |           |            |             |
|--------------|------------|-----------|------------|-------------|
| Atmosphäphil | Chalcophil | Lithophil | Siderophil | sehr selten |
|--------------|------------|-----------|------------|-------------|

# Radioaktiver Zerfall

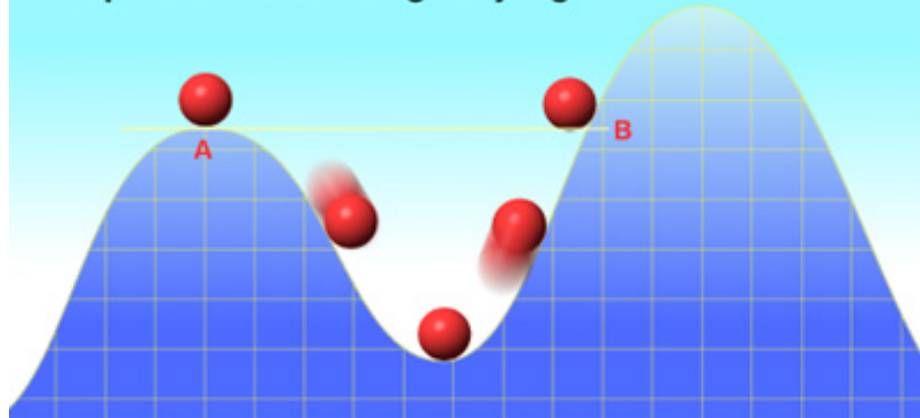
## Die (wichtigsten) Zerfallsarten



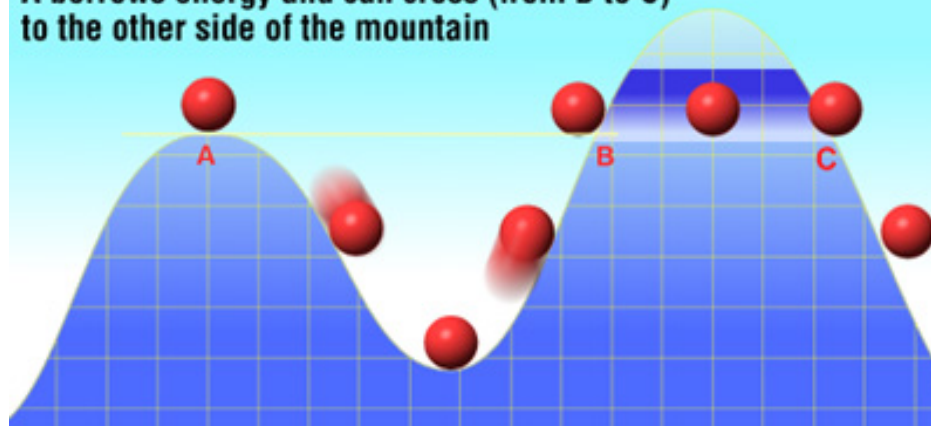
# Alpha ( $\alpha$ ) Zerfall

## Tunnel Effekt

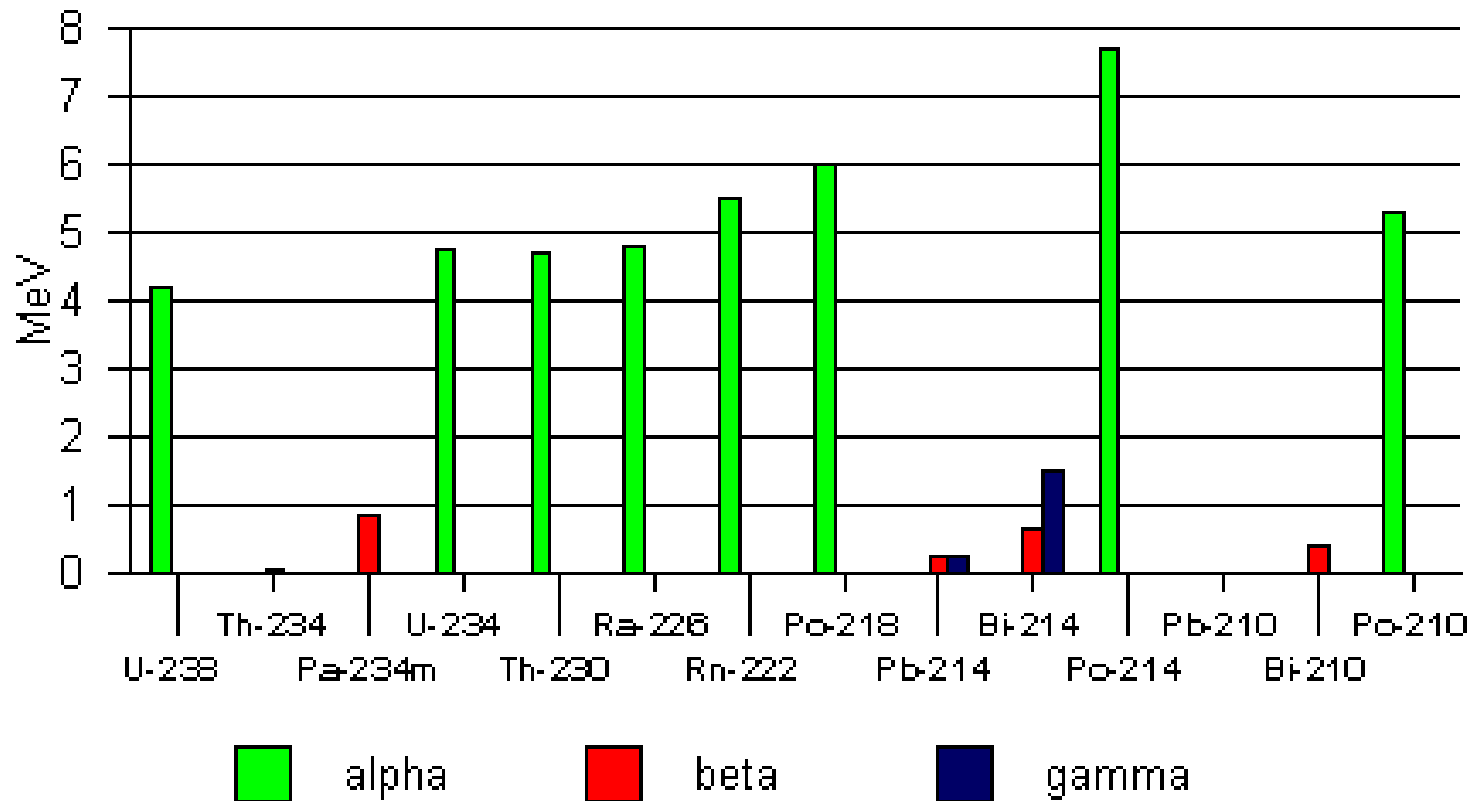
In classical dynamics,  
A stops at B and cannot go any higher



In quantum mechanics,  
A borrows energy and can cross (from B to C)  
to the other side of the mountain

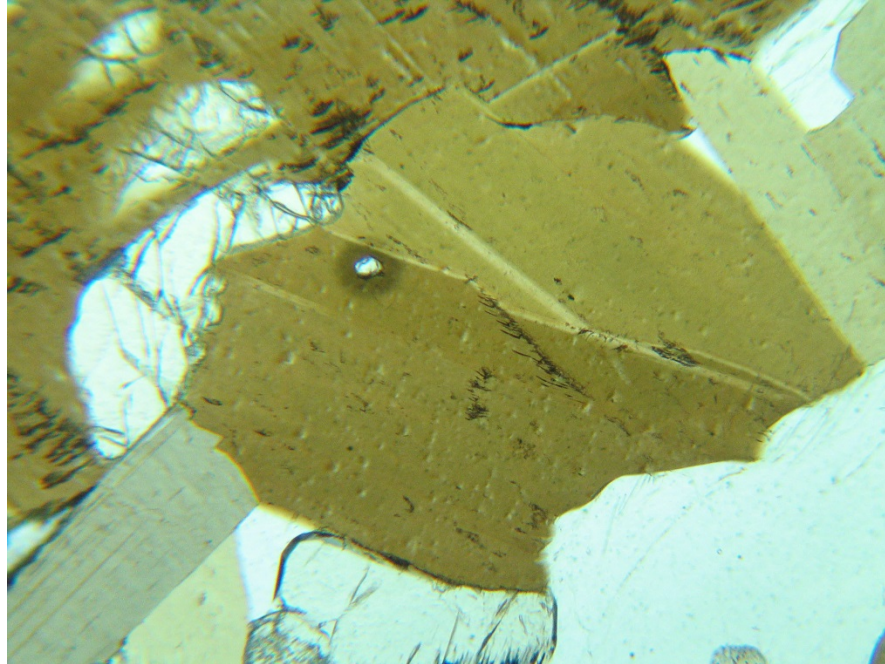


# Zerfallsenergien: U-238 Serie



# Alpha ( $\alpha$ ) Zerfall - Kristallgitterschäden

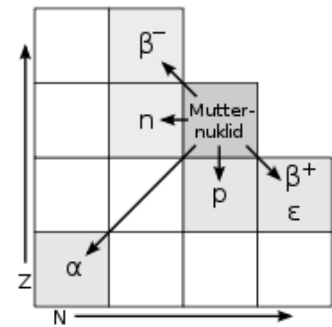
*Biotite with a halo around a zircon inclusion*



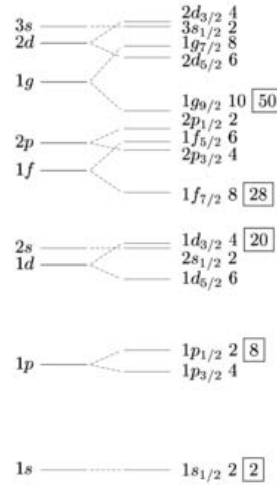
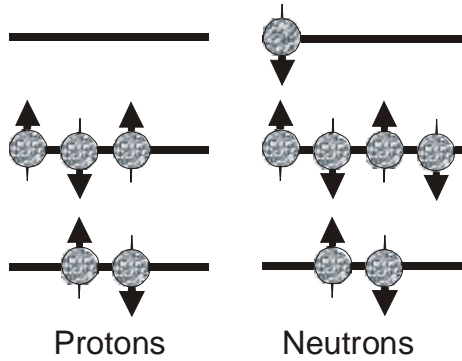
The field of view is about 2 mm

# $\beta$ -Zerfall

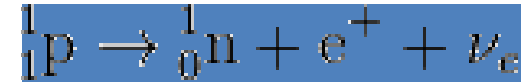
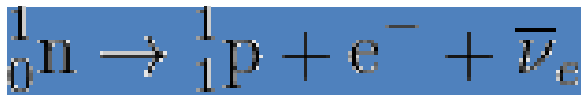
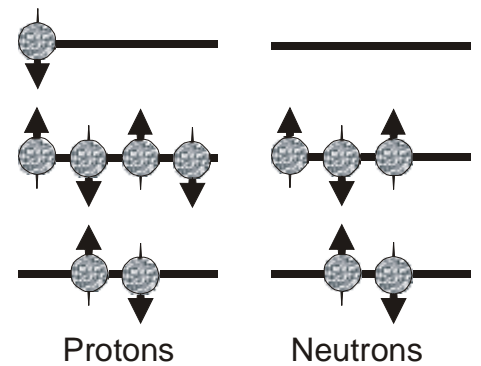
## Kernschalenmodell



$^{12}_5\text{B}$



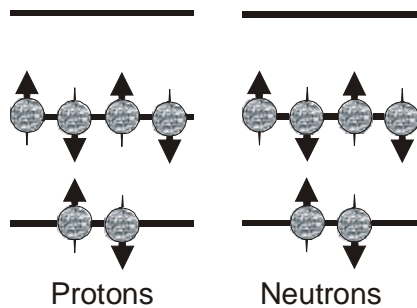
$^{12}_7\text{N}$



Beta Minus Decay



$^{12}_6\text{C}$



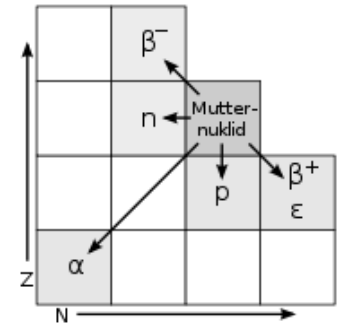
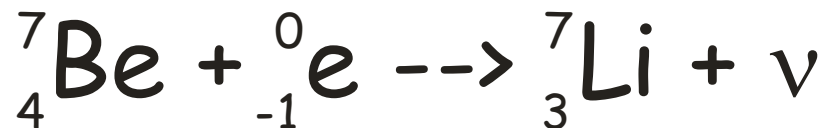
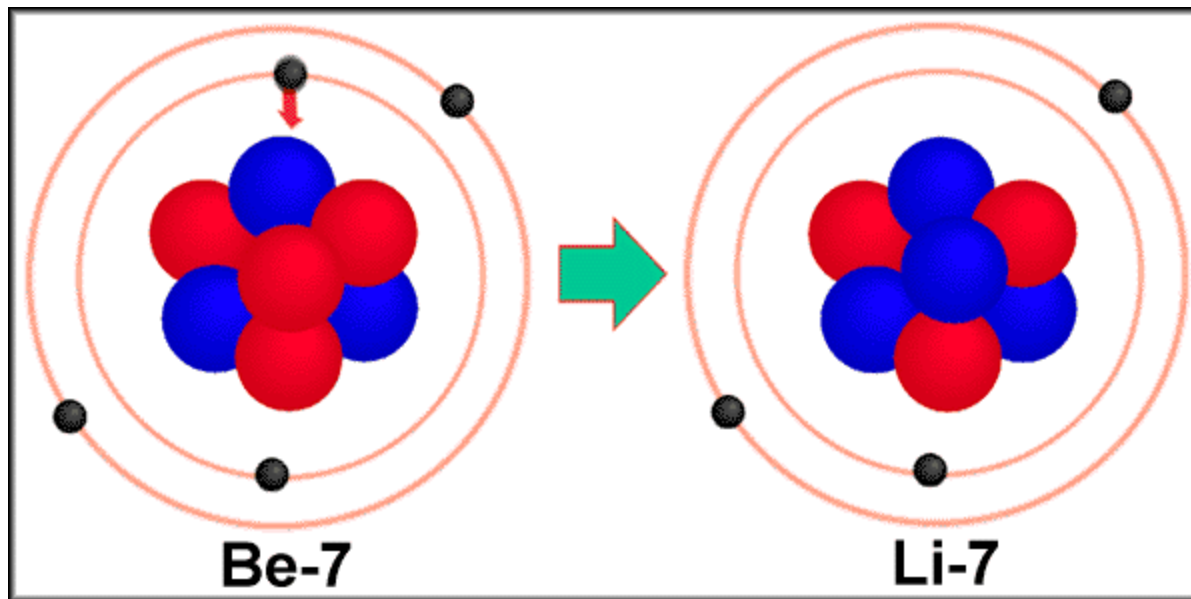
Beta Plus Decay



# Elektroneneinfang

A parent nucleus may capture one of its orbital electrons and emit a neutrino

Most commonly, it is a K-shell electron which is captured, and this is referred to as K-capture

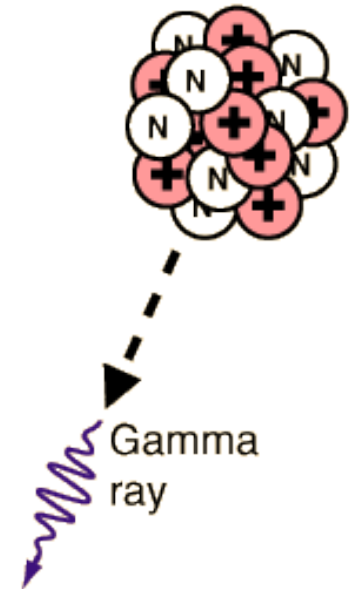
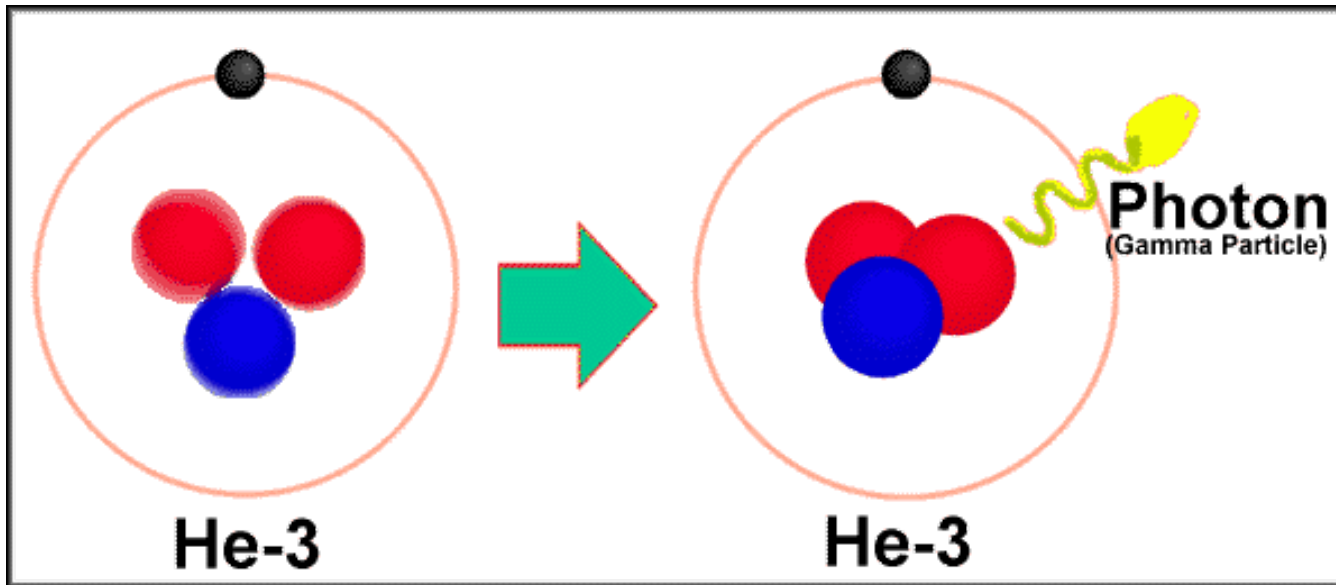




# Gamma( $\gamma$ ) Strahlung

Gamma radioactivity is composed of electromagnetic rays.

It is distinguished from x-rays by the fact that it comes from the nucleus.



He-3 produced by Tritium

# Elements

An International Magazine of Mineralogy, Geochemistry, and Petrology

February 2013  
Volume 9, Number 1

ISSN 1811-5209

## One Hundred Years of Geochronology

DANIEL J. CONDON and MARK D. SCHMITZ, Guest Editors

**...and Counting**

**Precision and Accuracy in Geochronology**

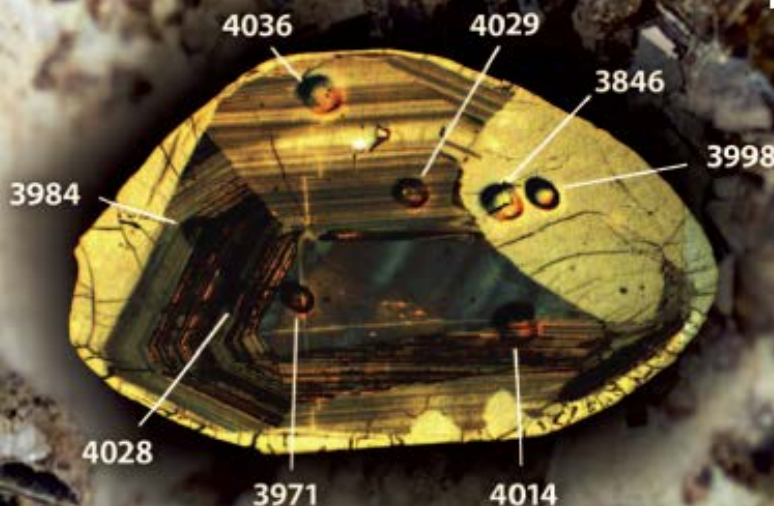
**High-Precision Geochronology**

**High-Spatial-Resolution Geochronology**

**Dating the Oldest Rocks  
in the Solar System**

**Time Constraints in the  
Quaternary Period**

**100 Years of U-Pb  
Geochronology**



Arthur Holmes 1913:  
*The age of the Earth*

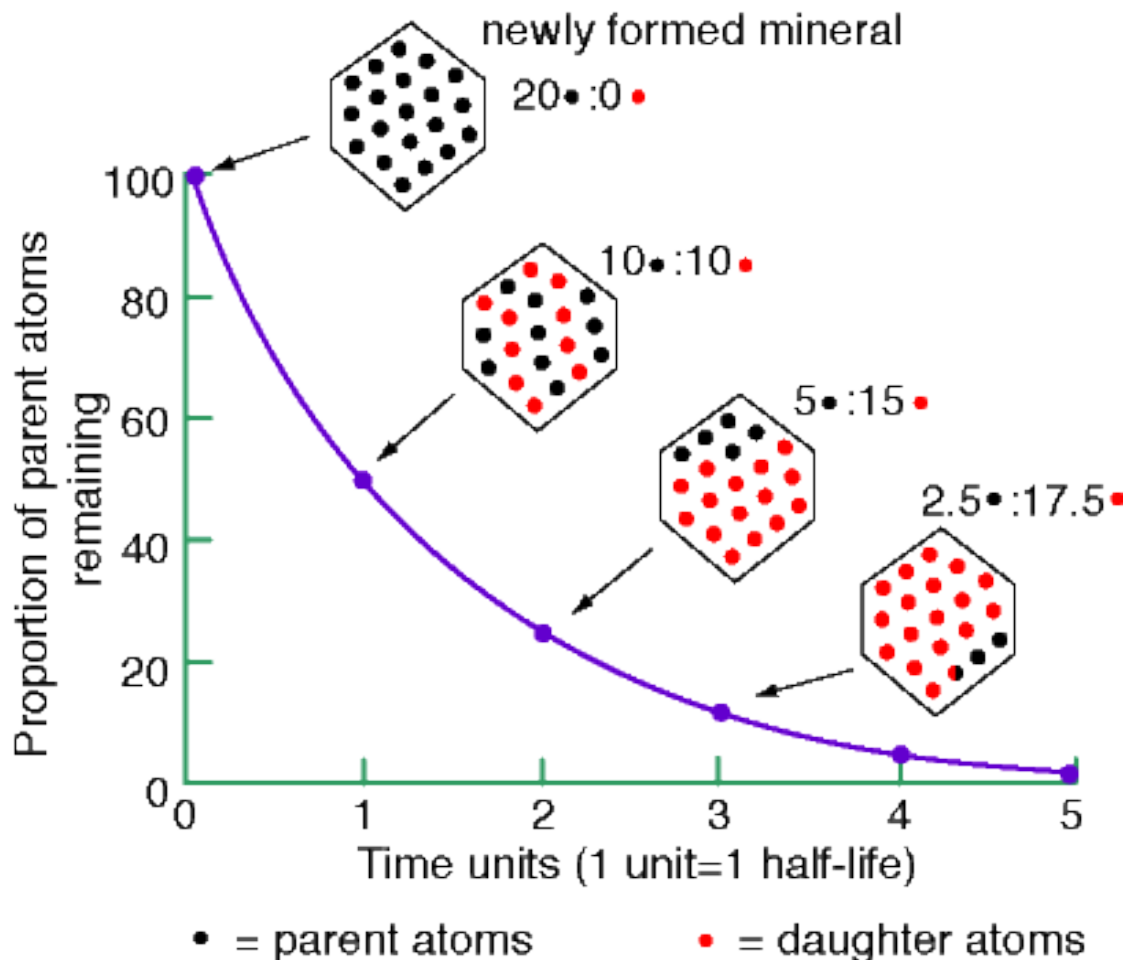
# Relative Ereignisabfolge?



Jutulhogget, Antarctica

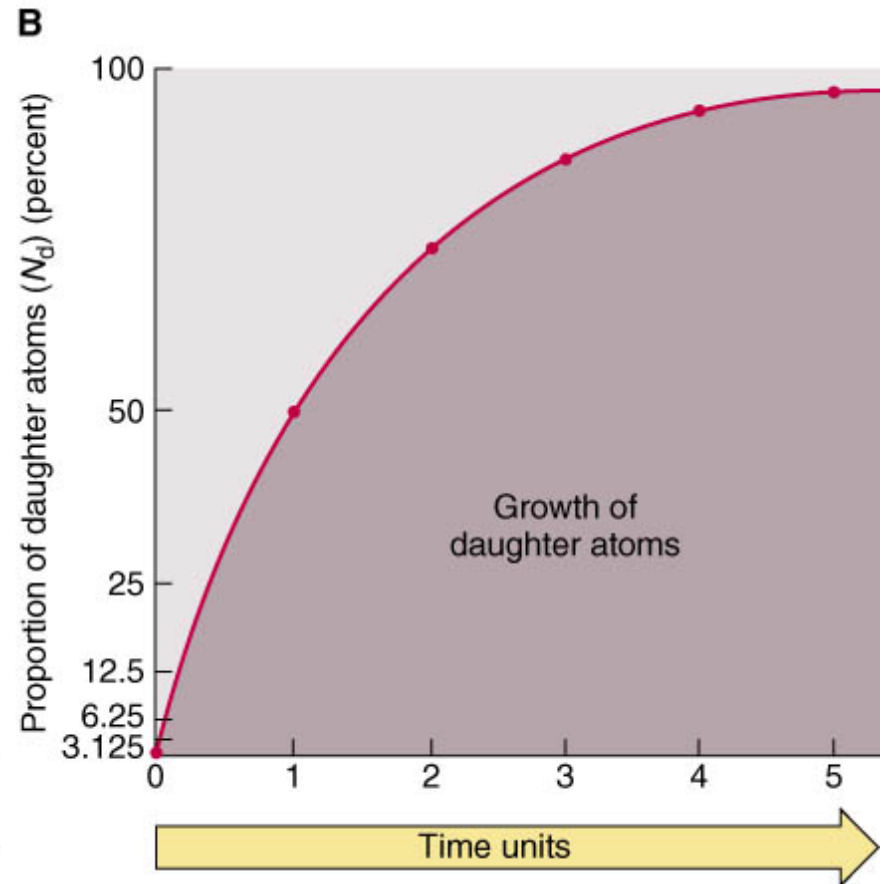
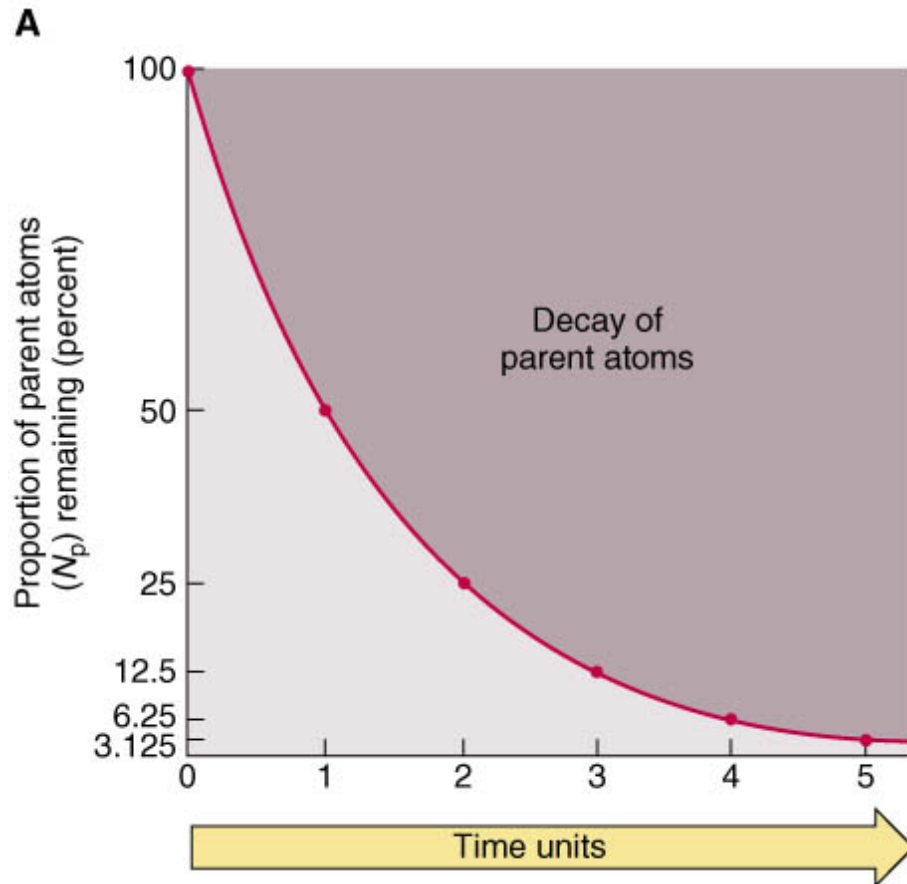
# Radioaktiver Zerfall

Halbwertszeiten,  $T_{1/2}$



- if it is possible to determine the ratio of the PARENT and DAUGHTER atoms, it is then possible to determine how long ago the decay process started → age determination

# Mutter – Tochter System



# Radioaktiver Zerfall

- Rate of decay is proportional to the number of decaying nuclei
- Integrate to find the change in N with time

$$\frac{dN}{dt} = -\lambda N$$

$$\frac{dN}{N} = -\lambda \cdot dt$$

$\lambda$  = decay constant

# Radioaktiver Zerfall

- Integrieren:

$$\int_{N_0}^{N(t)} \frac{1}{N} dN = - \int_0^t \lambda dt$$

- Resultat  $N(t)$ :

$$N = N_0 e^{-\lambda t}$$

Mutter-Tochter  
System:

$$D = D_0 + N(e^{\lambda t} - 1)$$

# Radioaktiver Zerfall

$$A = -\frac{dN}{dt} \quad \text{mit } A = \lambda \cdot N$$

$$-\lambda \cdot N = \frac{dN}{dt}$$

$$-\lambda \cdot dt = \frac{1}{N} \cdot dN$$

$$\int_0^t -\lambda \cdot dt' = \int_{N_0}^N \frac{1}{N'} \cdot dN'$$

$$-\lambda t - (-\lambda \cdot 0) = \ln(N) - \ln(N_0)$$

$$-\lambda t = \ln\left(\frac{N}{N_0}\right)$$

$$e^{-\lambda t} = \frac{N}{N_0}$$

$$N(t) = N_0 \cdot e^{-\lambda t}$$



# Radioaktiver Zerfall

- Integrieren:

$$\int_{N_0}^{N(t)} \frac{1}{N} dN = - \int_0^t \lambda dt$$

- Resultat  $N(t)$ :

$$N = N_0 e^{-\lambda t}$$

Mutter-Tochter  
System:

$$D = D_0 + N(e^{\lambda t} - 1)$$

# Radioaktiver Zerfall und Altersgleichung

Parent—daughter system:

$$D = N_0 - N$$

D – Number of daughter atoms, today

N – Number of parent atoms, today

$N_0$  – Number of parent atoms, initially present

$$N = N_0 \cdot e^{-\lambda \cdot t} \quad \text{or} \quad N_0 = N \frac{1}{e^{-\lambda t}} = N e^{\lambda t}$$

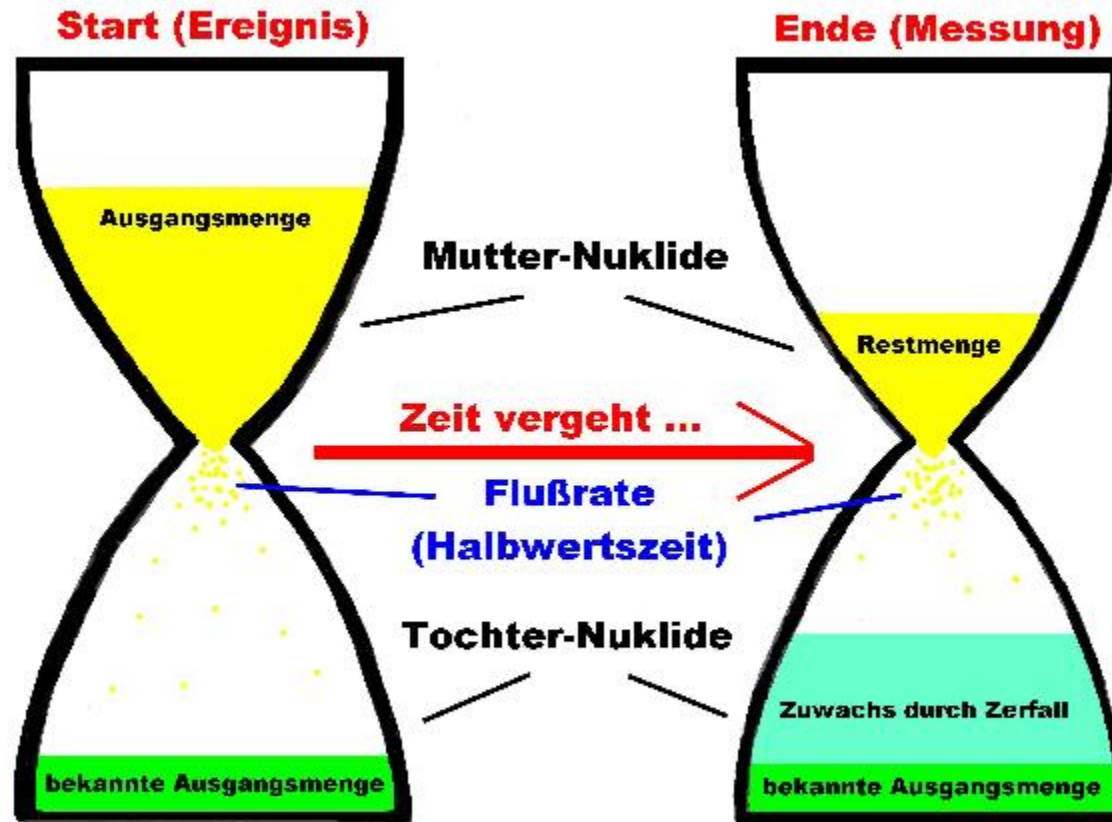
$$N_0 = D + N, \text{ hence: } D + N = N e^{\lambda t}, \text{ or } D = N e^{\lambda t} - N, \text{ or } D = N(e^{\lambda t} - 1)$$

$$t = \frac{1}{\lambda} \ln \left( 1 + \frac{D}{N} \right)$$

**This is the mathematical expression that relates radioactive decay to geologic time!**

If some daughter nuclides,  $D_0$ , are there initially:  **$D = D_0 + N(e^{\lambda t} - 1)$**

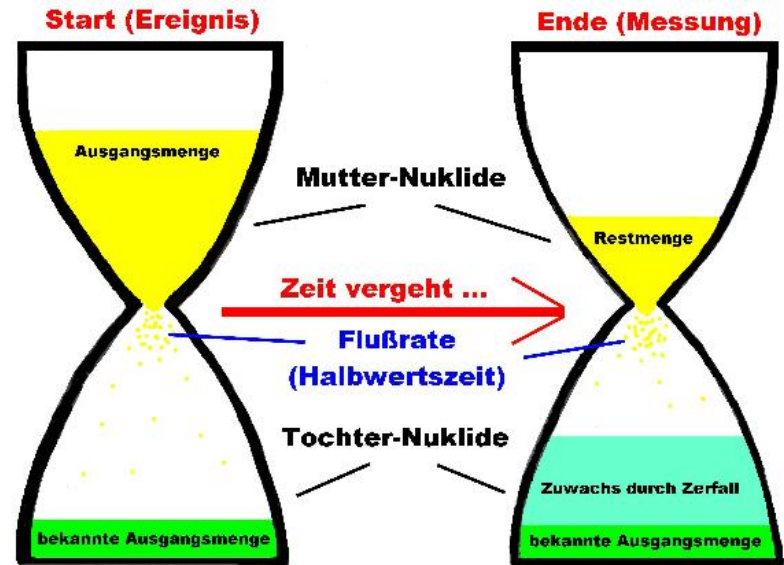
# Radiometrische Datierung



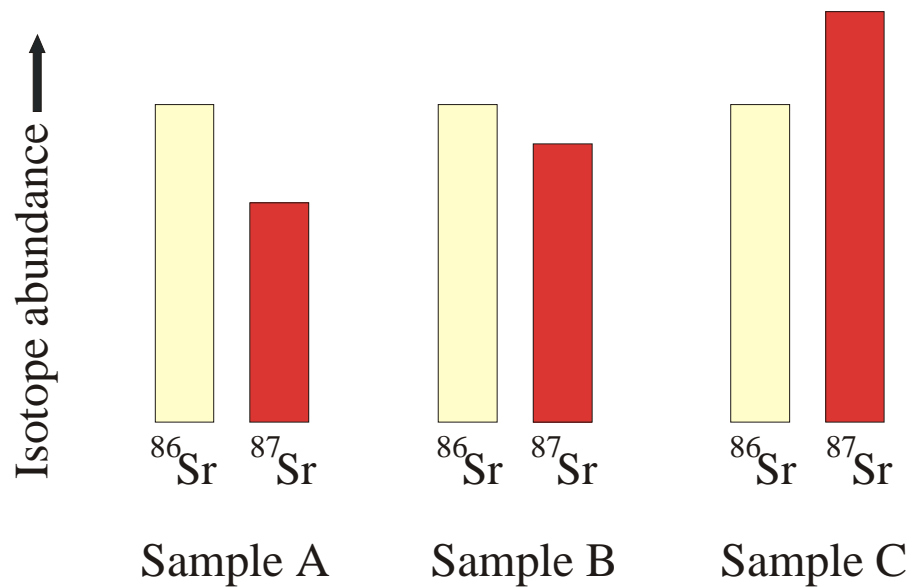
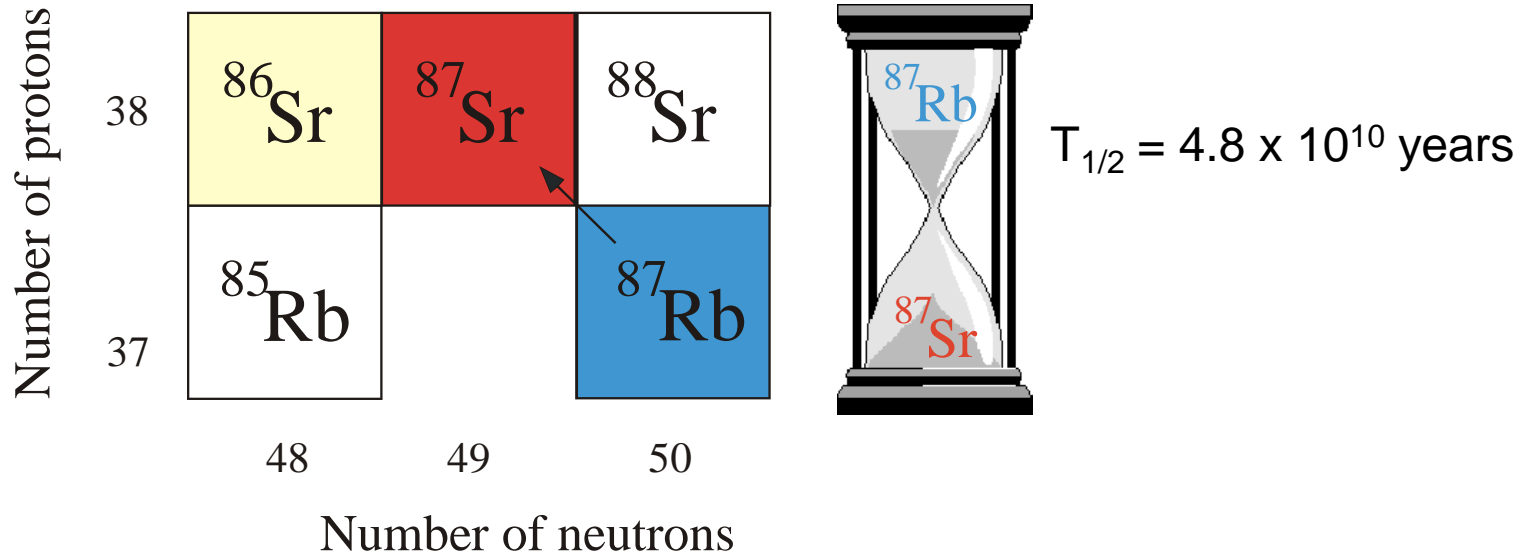
# Radiometrische Datierung

Method is restricted to minerals which:

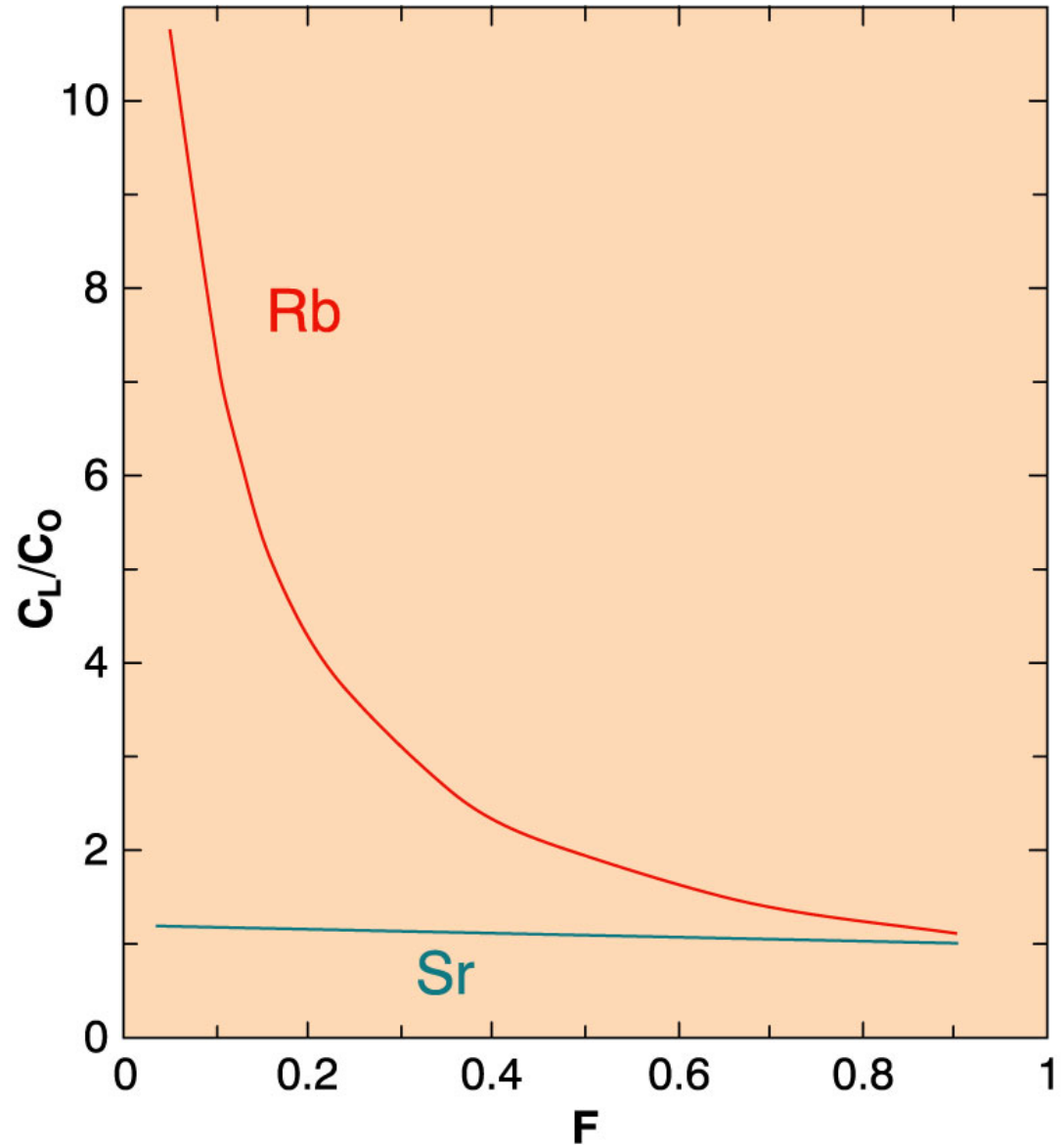
- Still contain some of the parent nuclei
- Allowed for no gain or loss of D or P as time passed
- Initially contained no D (or  $D_0$  must be known)



# Rb–Sr System

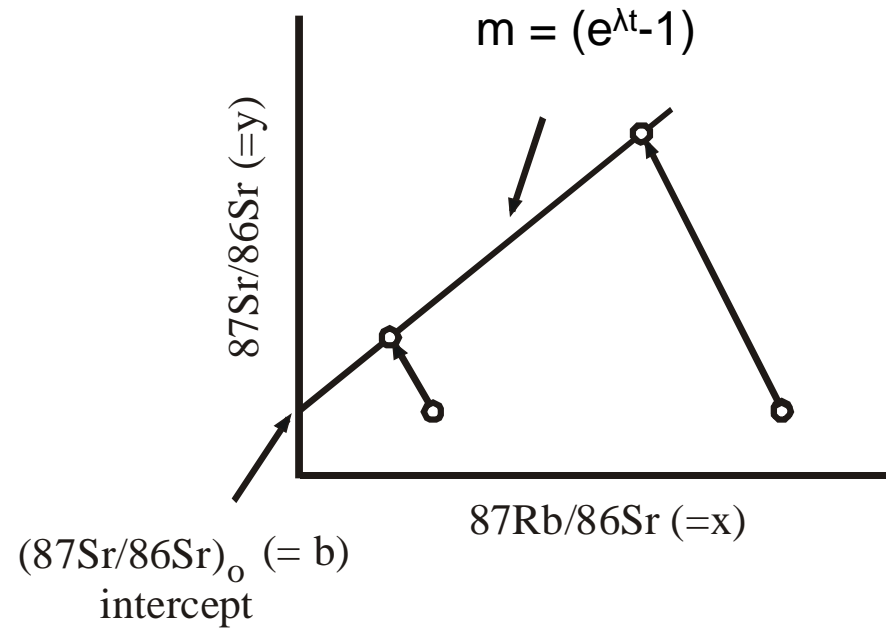
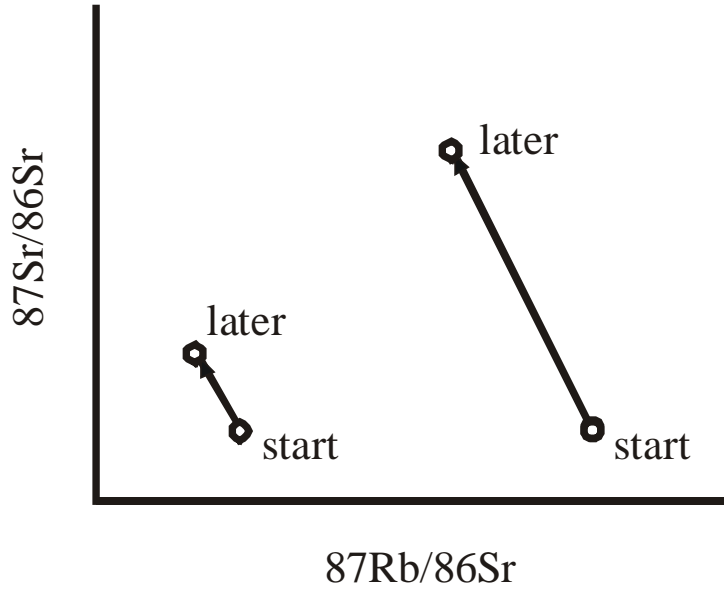


# Geochemische Diagramme



Veränderung der Rb und Sr  
Konzentrationen in der Schmelze mit  
steigendem Aufschmelzungsgrad.  
Basaltisches Ausgangsgestein  
(Plagioklas, Augit, Olivin).

# Rb–Sr System – Nicolaysen Diagramm



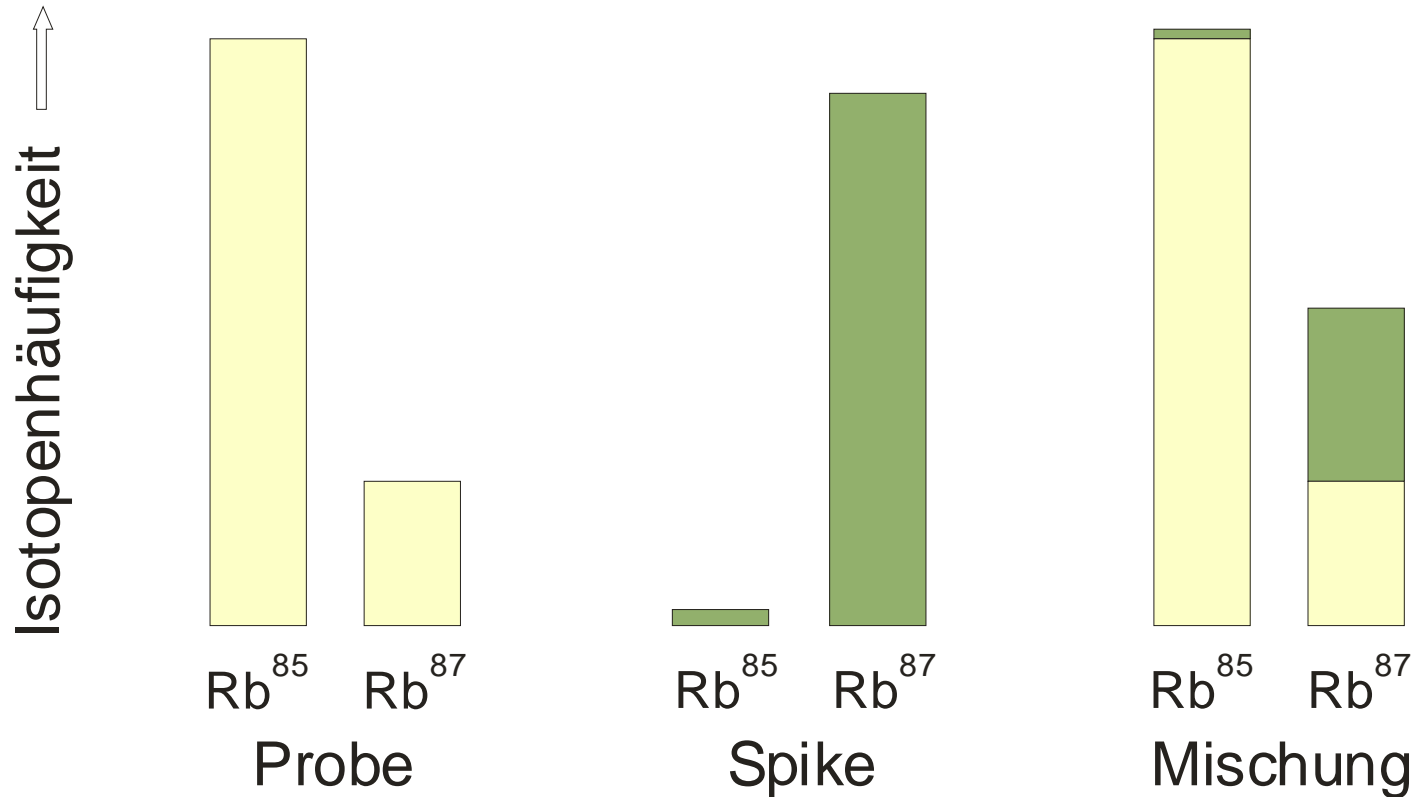
$${}^{87}\text{Sr}/{}^{86}\text{Sr} = ({}^{87}\text{Sr}/{}^{86}\text{Sr})_0 + {}^{87}\text{Rb}/{}^{86}\text{Sr} (e^{\lambda t}-1)$$

$$y = b + x m$$

$$m = (e^{\lambda t}-1)$$

$$T = \ln(m + 1)/\lambda$$

# Isotopenverdünnungsanalyse

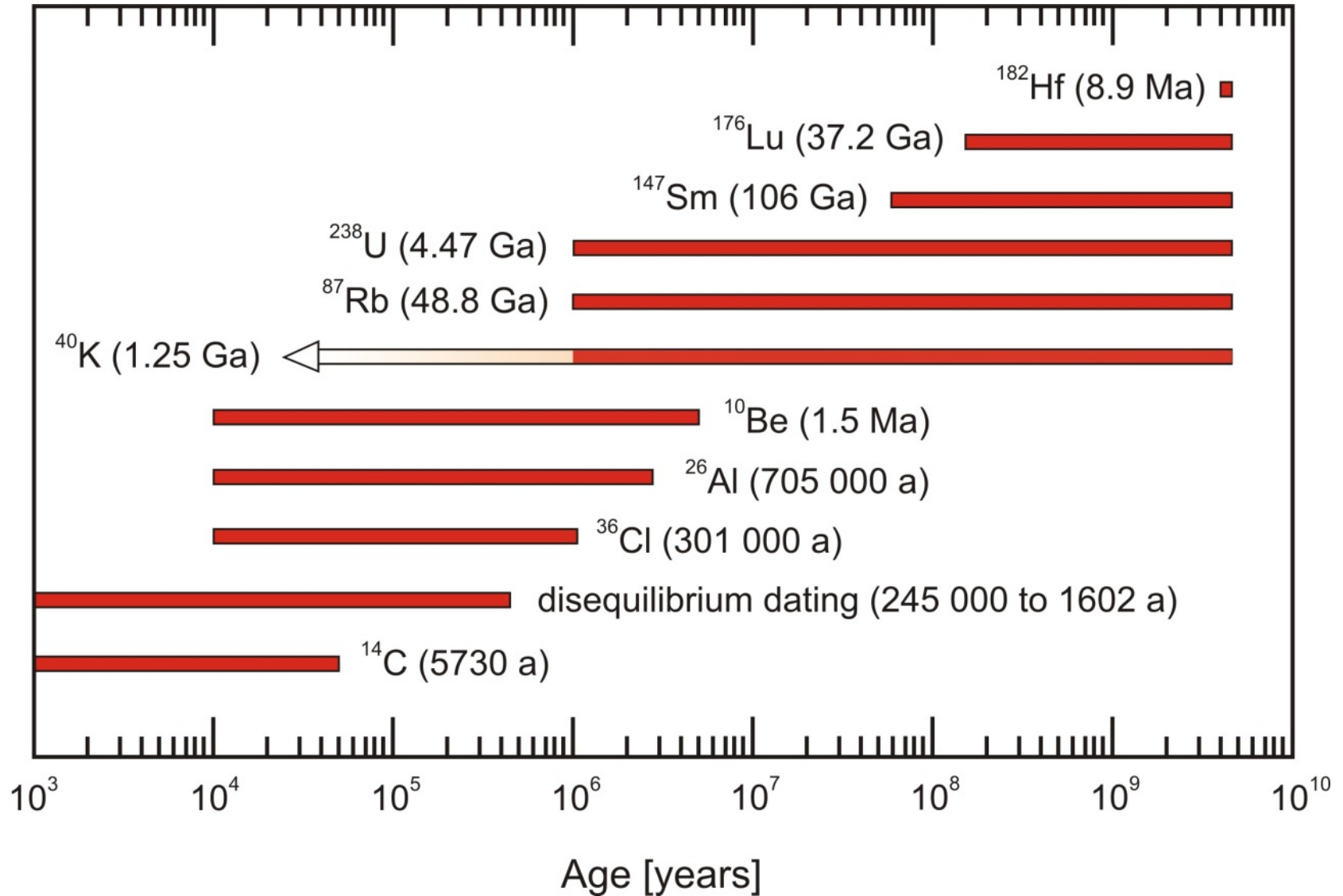




# Isotopenverdünnungsanalyse

- Zusatz des gleichen Elements mit anderer Isotopenzusammensetzung (Spike)
- Mischung von Probe und Isotopenstandard
- Abtrennung des Elements (keine quantitative Ausbeute notwendig!)
- Bestimmung des Verhältnis der Isotope
- Berechnung der Konzentration

# Datierungssysteme

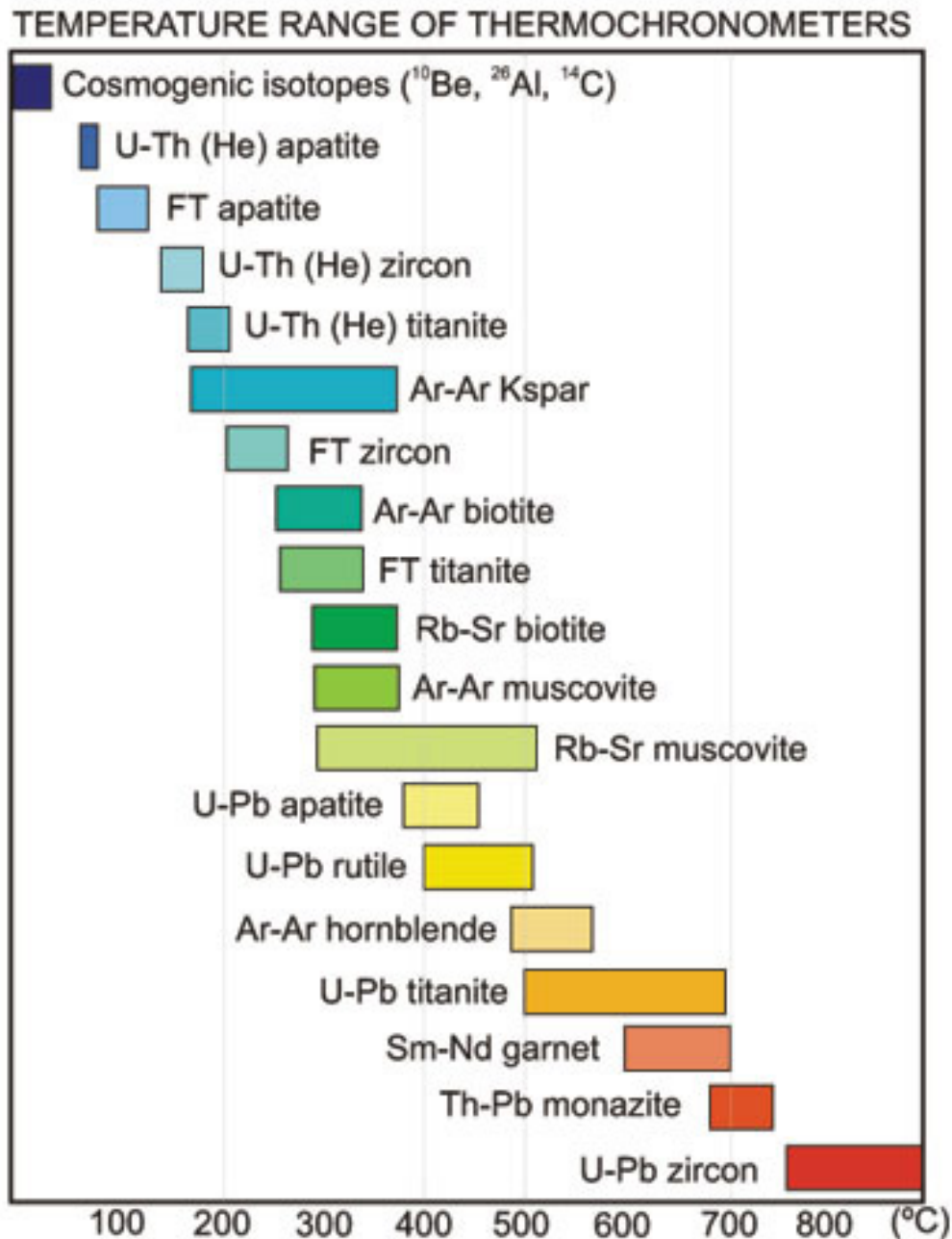


# Datierungsmethoden und Schließungstemperatur

Schließungstemperatur: Temperatur unterhalb der keine signifikante Diffusion mehr stattfindet und die geochronologische Uhr zu ticken beginnt

| <b>Mineral</b>    | <b>Method</b> | <b>T (°C)</b> |
|-------------------|---------------|---------------|
| Zircon            | U-Pb          | >800          |
| Monazite          | U-Pb          | >800          |
| Titanite (Sphene) | U-Pb          | 600           |
| Garnet            | Sm-Nd         | >550          |
| Hornblende        | K-Ar          | 500           |
| Muscovite         | Rb-Sr         | 500           |
| Muscovite         | K-Ar          | 350           |
| Apatite           | U-Pb          | 350           |
| Biotite           | Rb-Sr         | 300           |
| Biotite           | K-Ar          | 280           |
| K-Feldspar        | K-Ar          | 200           |
| Apatite           | Fission Track | 120           |

# Datierungsmethoden und Schließungstemperatur



# Welche Art von Alter?

**Bildungsalter** (Entstehungsalter)

oder ererbtes Alter?

**Abkühlungsalter** (Abkühlalter) oder

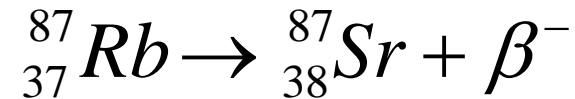
Metamorphosealter (Aufheizungsalter)?

**Hebungsalter** (bei Spaltspuren)

Welches Mineral oder Gestein wurde mit welcher Methode datiert?

# Isochronenmethode

Beispiel Rb-Sr:



Halbwertszeit:  $T_{1/2} = 48.8 \text{ Ga}$

Zerfallskonstante:  $\lambda = \ln 2 / T_{1/2} = 0.693 / T_{1/2}$

Altersgleichung:

$$\frac{{}^{87}\text{Sr}}{{}^{86}\text{Sr}} = \left( \frac{{}^{87}\text{Sr}}{{}^{86}\text{Sr}} \right)_0 + \frac{{}^{87}\text{Rb}}{{}^{86}\text{Sr}} (e^{\lambda t} - 1)$$

# Rb-Sr Isochronendiagramm

$M_1, M_2$  = cogenetische Minerale

$R_1, R_2$  cogenetische Gesamtgesteine mit unterschiedlichen Rb/Sr Verhältnissen



Margin

Interior



Tonalite

to

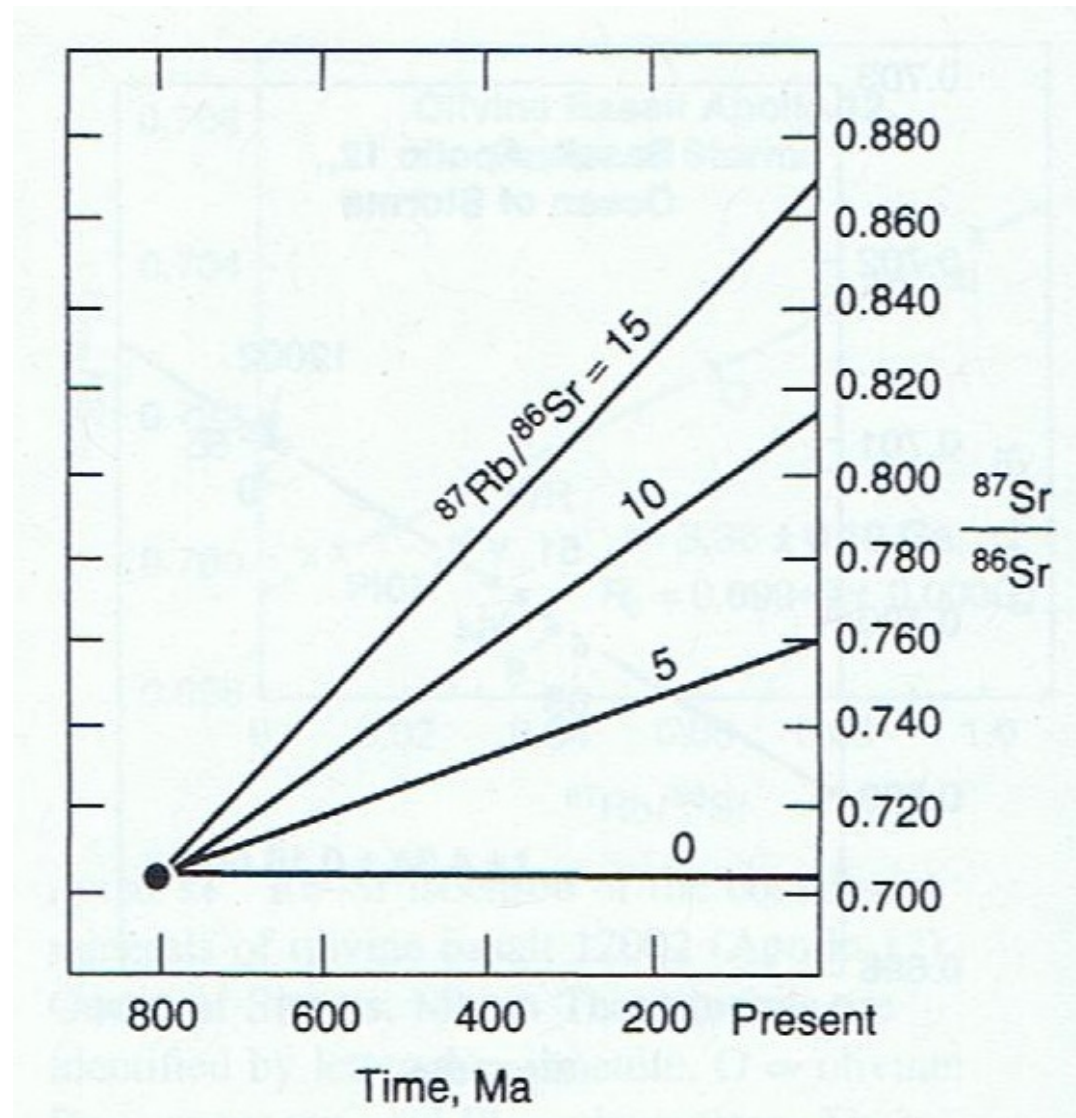
Granodiorite

Monzonite

to

Granite

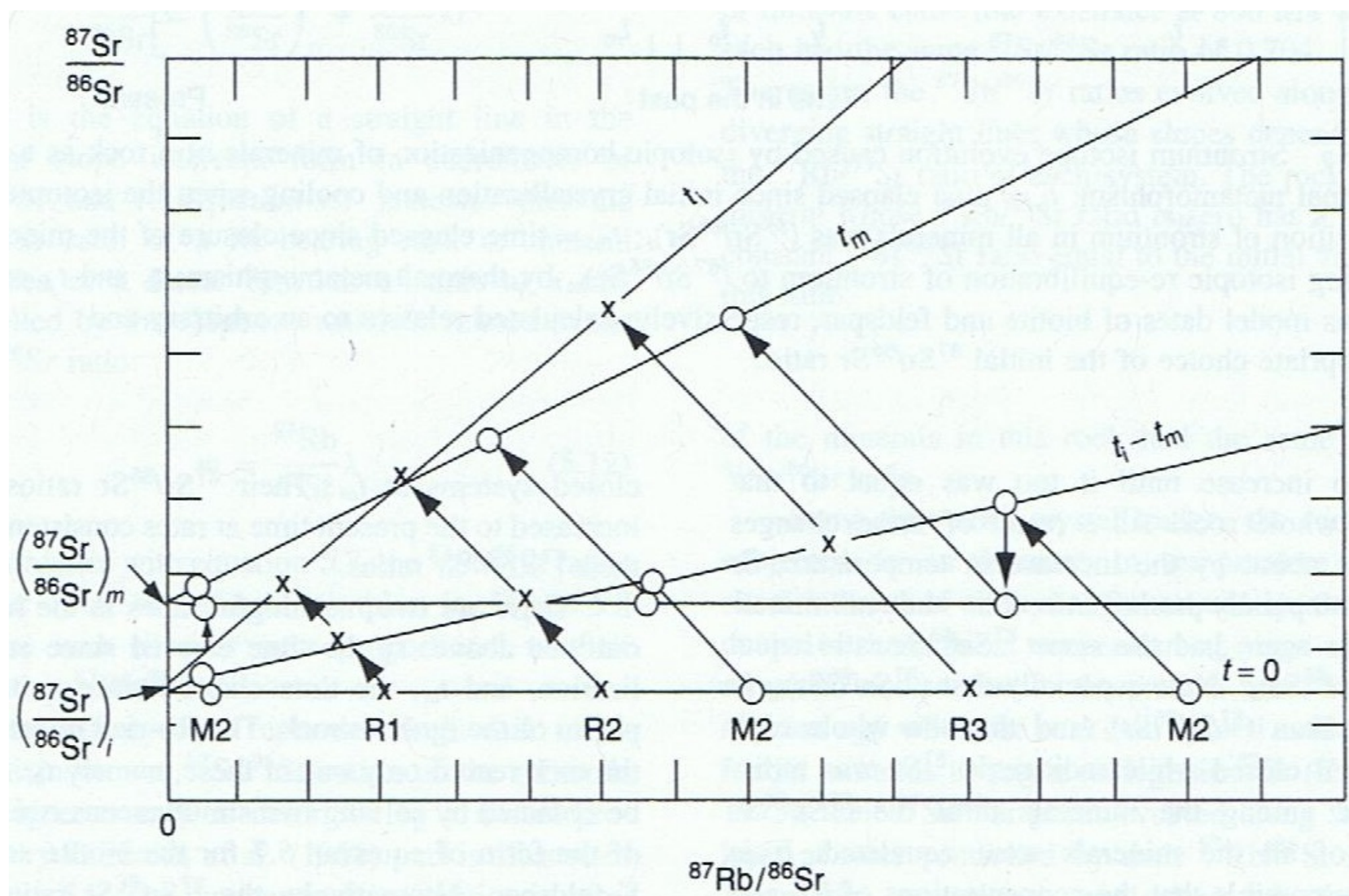
# Compston-Jeffery Diagramm



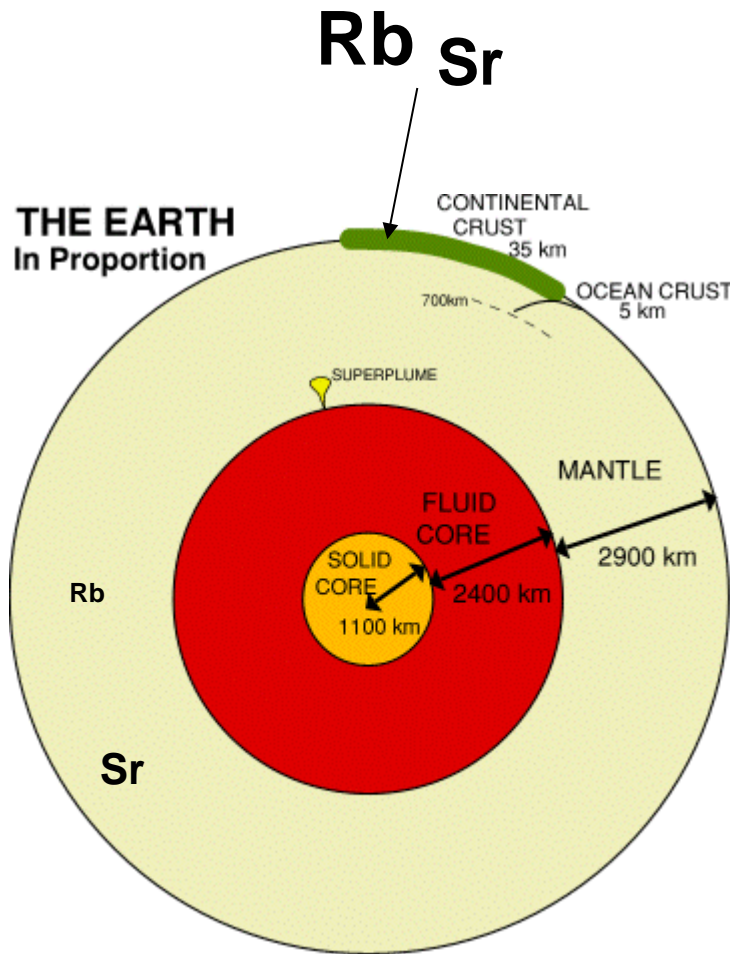




# Rb/Sr-System während einer Metamorphose

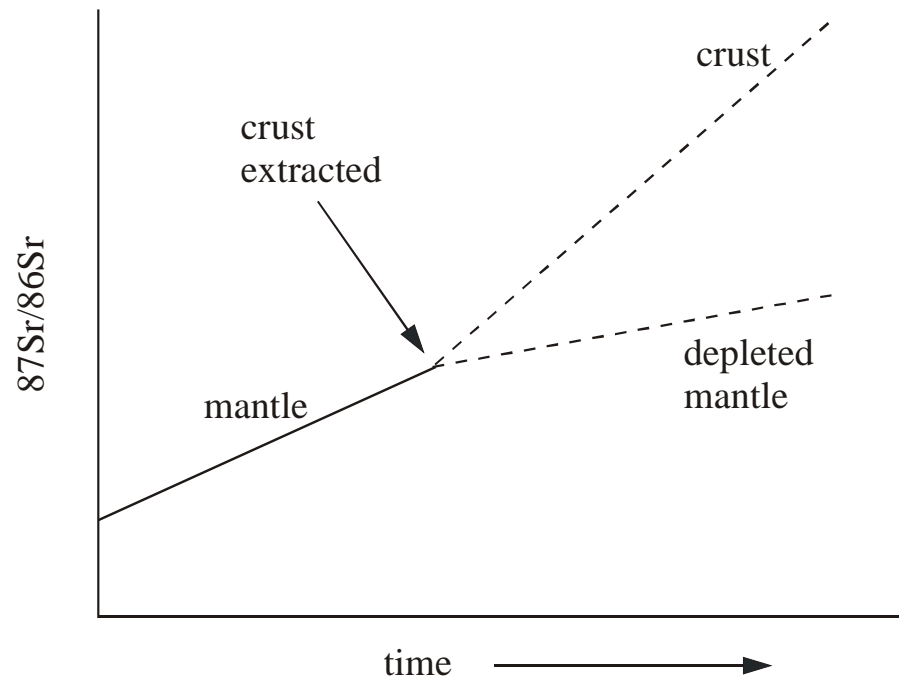


# Die Sr-Isotopenentwicklung der Erde



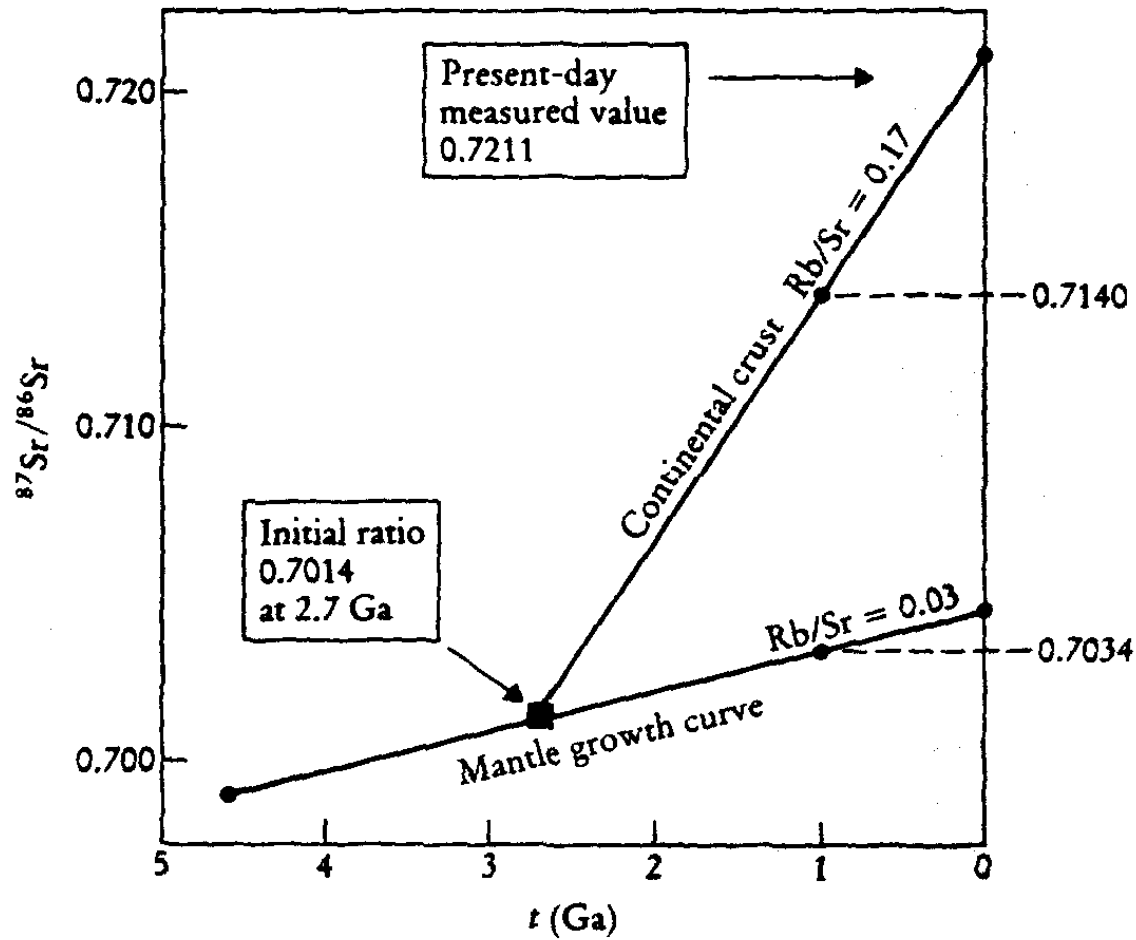
Mehr Rb in der Kruste als im Erdmantel

Dadurch höhere  $^{87}\text{Sr}/^{86}\text{Sr}$  Verhältnisse in der Kruste als im Mantel



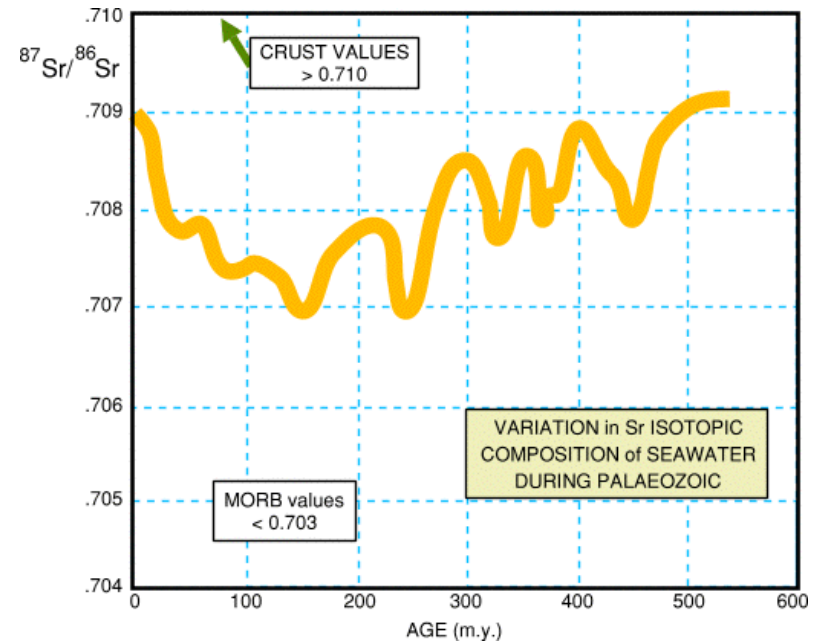
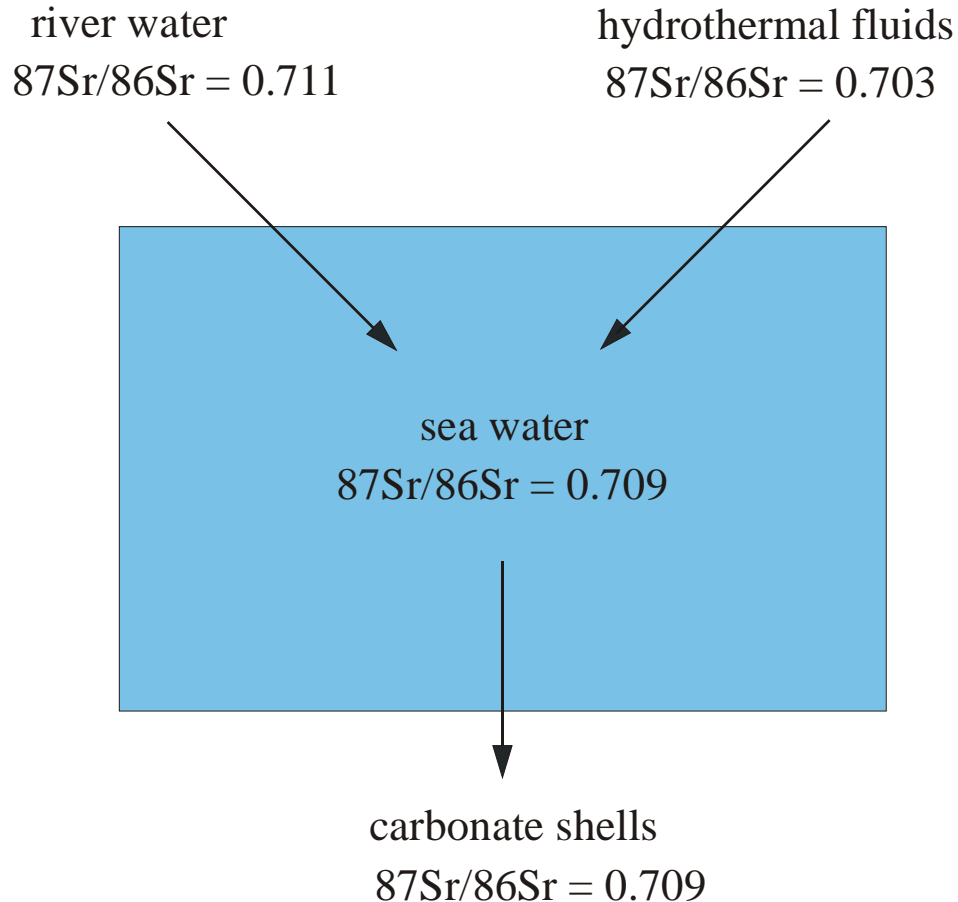
Continental crust: 32-78 ppm Rb, 260-333 ppm Sr  
Depleted Mantle: 0.6 ppm Rb, 19.9 ppm Sr

# $^{87}\text{Sr}/^{86}\text{Sr}$ Isotopenentwicklung in Kruste und Mantel



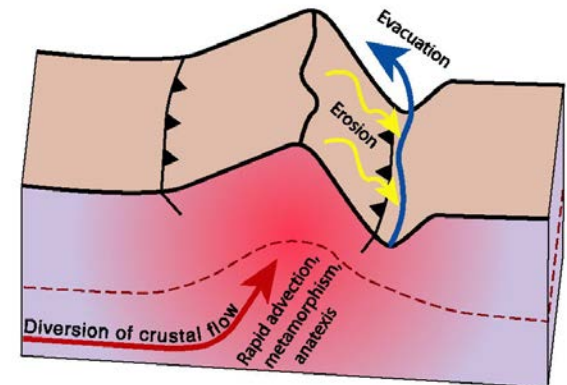
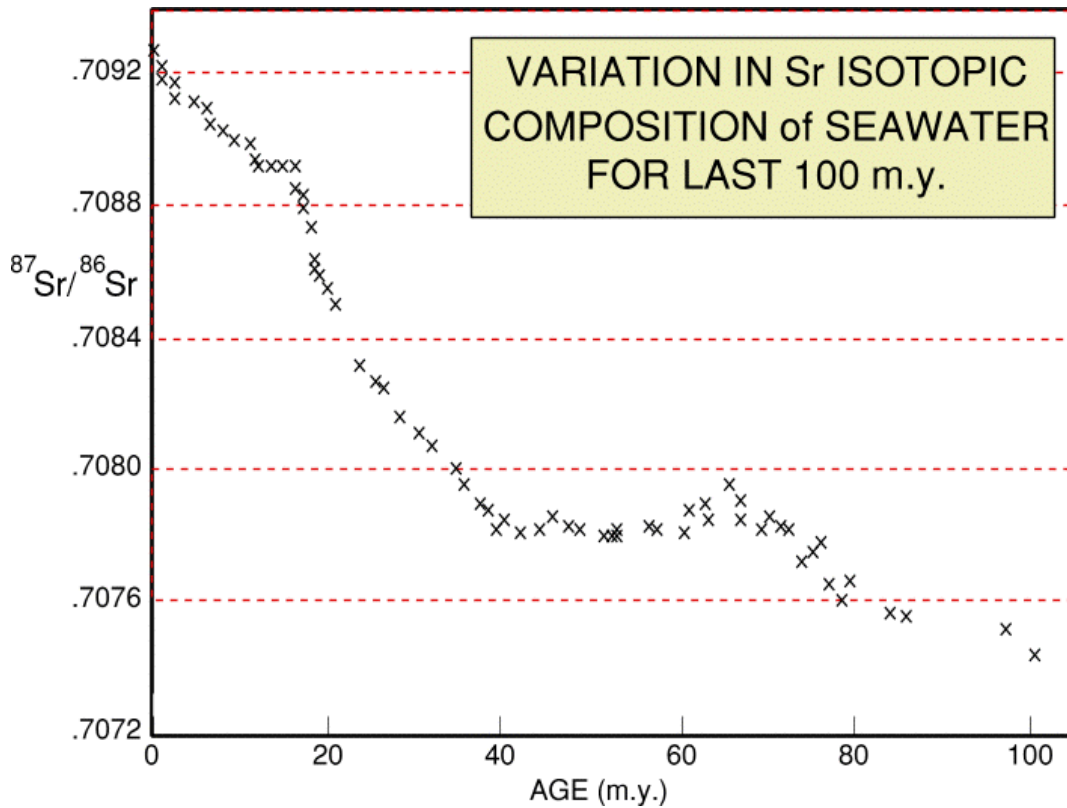
# Sr-Isotopie der Ozeane

Sr isotope composition of the oceans is determined by the relative contributions of Sr from river waters and hydrothermal sources



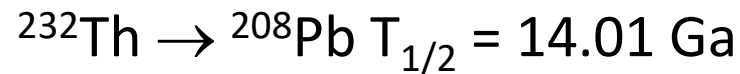
# Sr-Isotopie der Ozeane

Increase in the global ocean  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio since India-Asia collision



# Die U-Th-Pb Methoden

Basiert auf drei verschiedenen Zerfallsreaktionen:



$^{204}\text{Pb}$  ist ein nicht radiogen, stabiles Isotop. Daher gilt:

$$\frac{^{206}\text{Pb}}{^{204}\text{Pb}} = \left( \frac{^{206}\text{Pb}}{^{204}\text{Pb}} \right)_0 + \frac{^{238}\text{U}}{^{204}\text{Pb}} (e^{\lambda_1 t} - 1)$$

$$\frac{^{207}\text{Pb}}{^{204}\text{Pb}} = \left( \frac{^{207}\text{Pb}}{^{204}\text{Pb}} \right)_0 + \frac{^{235}\text{U}}{^{204}\text{Pb}} (e^{\lambda_2 t} - 1)$$

$$\frac{^{208}\text{Pb}}{^{204}\text{Pb}} = \left( \frac{^{208}\text{Pb}}{^{204}\text{Pb}} \right)_0 + \frac{^{232}\text{Th}}{^{204}\text{Pb}} (e^{\lambda_3 t} - 1)$$

# Die U-Th-Pb Methoden

Diese drei Gleichungen liefern drei voneinander unabhängige Alter von Th und U-haltigen Gesteinen oder Mineralen

Wenn kein Austausch von U, Th oder Pb stattgefunden hat (geschlossenes System) sollten diese Alter identisch sein

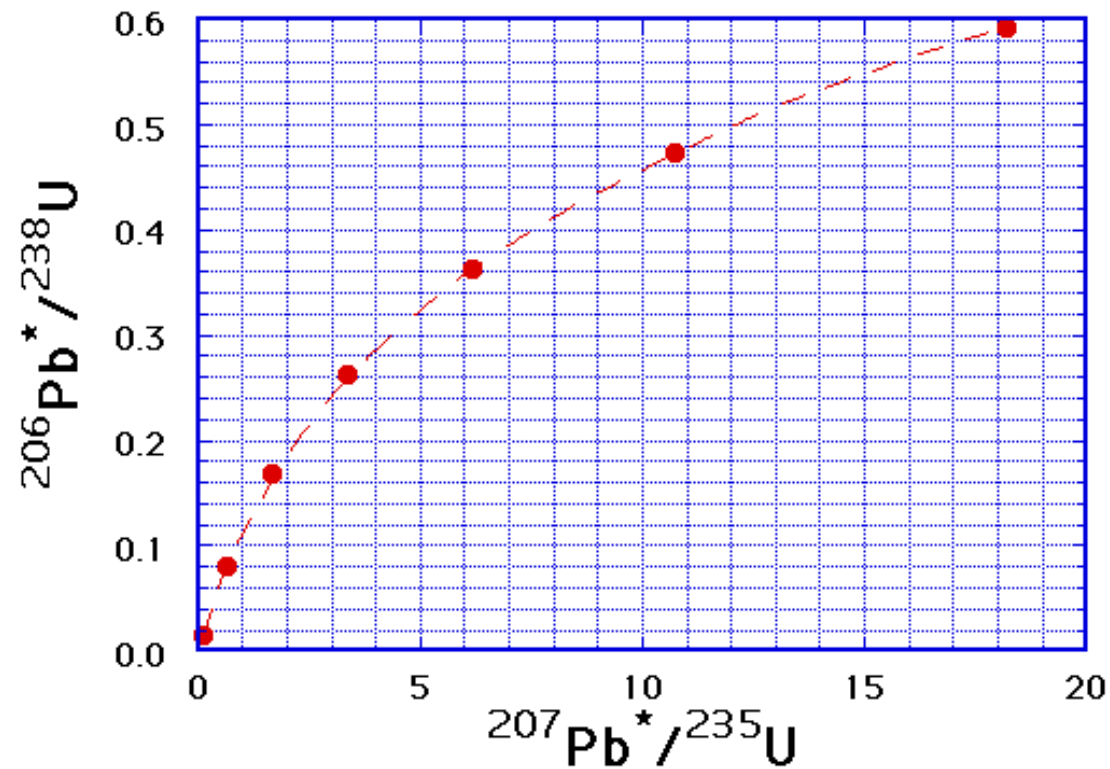
Wenn die drei Alter übereinstimmen, sind sie **konkordant**, sind sie unterschiedlich sind, sind sie **diskordant**

Eine häufige Ursache für diskordante Altersdaten ist **Bleiverlust**



# U-Pb Concordia Diagram

$$^{206}\text{Pb}^*/^{238}\text{U} = (e^{\lambda t_1} - 1), \quad ^{207}\text{Pb}^*/^{235}\text{U} = (e^{\lambda t_2} - 1)$$



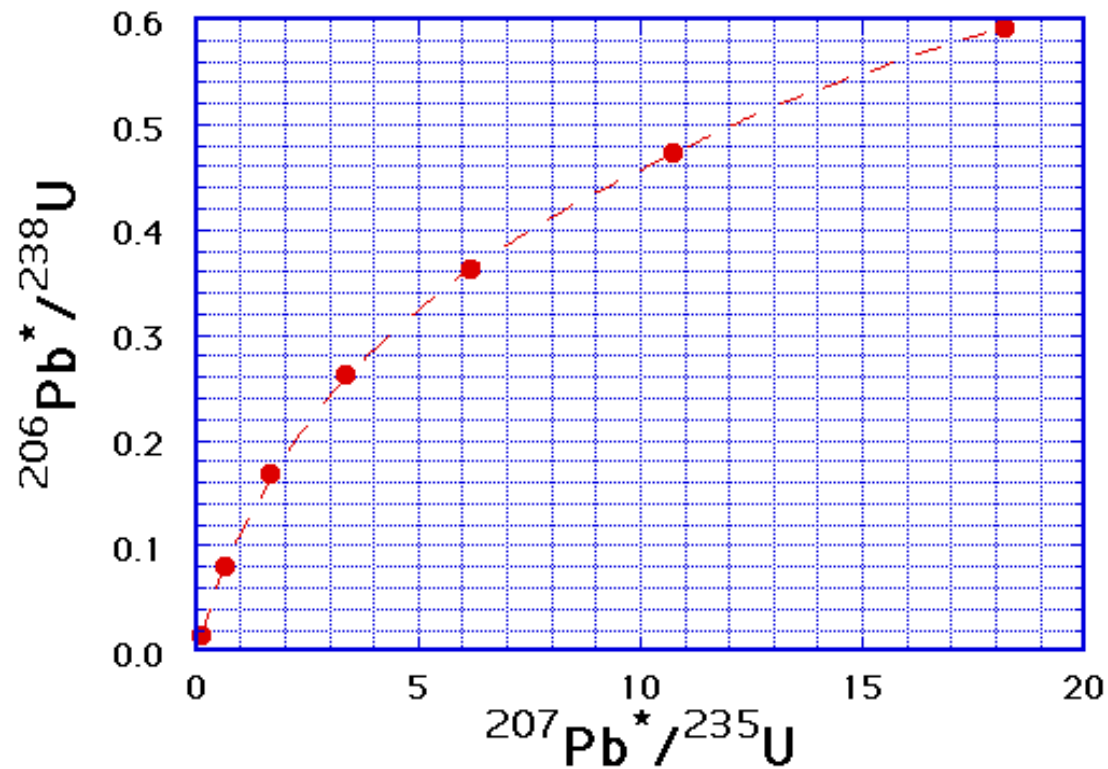
# U-Pb Concordia Diagramm

Numerical values of  $e^{\lambda_1 t} - 1$ ,  $e^{\lambda_2 t} - 1$ , and of the radiogenic  $^{207}\text{Pb}/^{206}\text{Pb}$  ratio as a function of age (t)

| Ga  | $e^{\lambda_1 t} - 1$ | $e^{\lambda_2 t} - 1$ | $\left(\frac{^{207}\text{Pb}}{^{206}\text{Pb}}\right)^*$ | Ga  | $e^{\lambda_1 t} - 1$ | $e^{\lambda_2 t} - 1$ | $\left(\frac{^{207}\text{Pb}}{^{206}\text{Pb}}\right)^*$ |
|-----|-----------------------|-----------------------|--|-----|-----------------------|-----------------------|--|
| 0   | 0.0000                | 0.0000                | 0.04607  | 2.4 | 0.4511                | 9.6296                | 0.15492  |
| 0.2 | 0.0315                | 0.2177                | 0.05014  | 2.6 | 0.4968                | 11.9437               | 0.17447  |
| 0.4 | 0.0640                | 0.4828                | 0.05473  | 2.8 | 0.5440                | 14.7617               | 0.19693  |
| 0.6 | 0.0975                | 0.8056                | 0.05994  | 3.0 | 0.5926                | 18.1931               | 0.22279  |
| 0.8 | 0.1321                | 1.1987                | 0.06584  | 3.2 | 0.6428                | 22.3716               | 0.25257  |
| 1.0 | 0.1678                | 1.6774                | 0.07254  | 3.4 | 0.6946                | 27.4597               | 0.28690  |
| 1.2 | 0.2046                | 2.2603                | 0.08017  | 3.6 | 0.7480                | 33.6556               | 0.32653  |
| 1.4 | 0.2426                | 2.9701                | 0.08886  | 3.8 | 0.8030                | 41.2004               | 0.37232  |
| 1.6 | 0.2817                | 3.8344                | 0.09877  | 4.0 | 0.8599                | 50.3878               | 0.42525  |
| 1.8 | 0.3221                | 4.8869                | 0.11010  | 4.2 | 0.9185                | 63.5753               | 0.48951  |
| 2.0 | 0.3638                | 6.1685                | 0.12306  | 4.4 | 0.9789                | 75.1984               | 0.55746  |
| 2.2 | 0.4067                | 7.7291                | 0.13790  | 4.6 | 1.0413                | 91.7873               | 0.63969  |

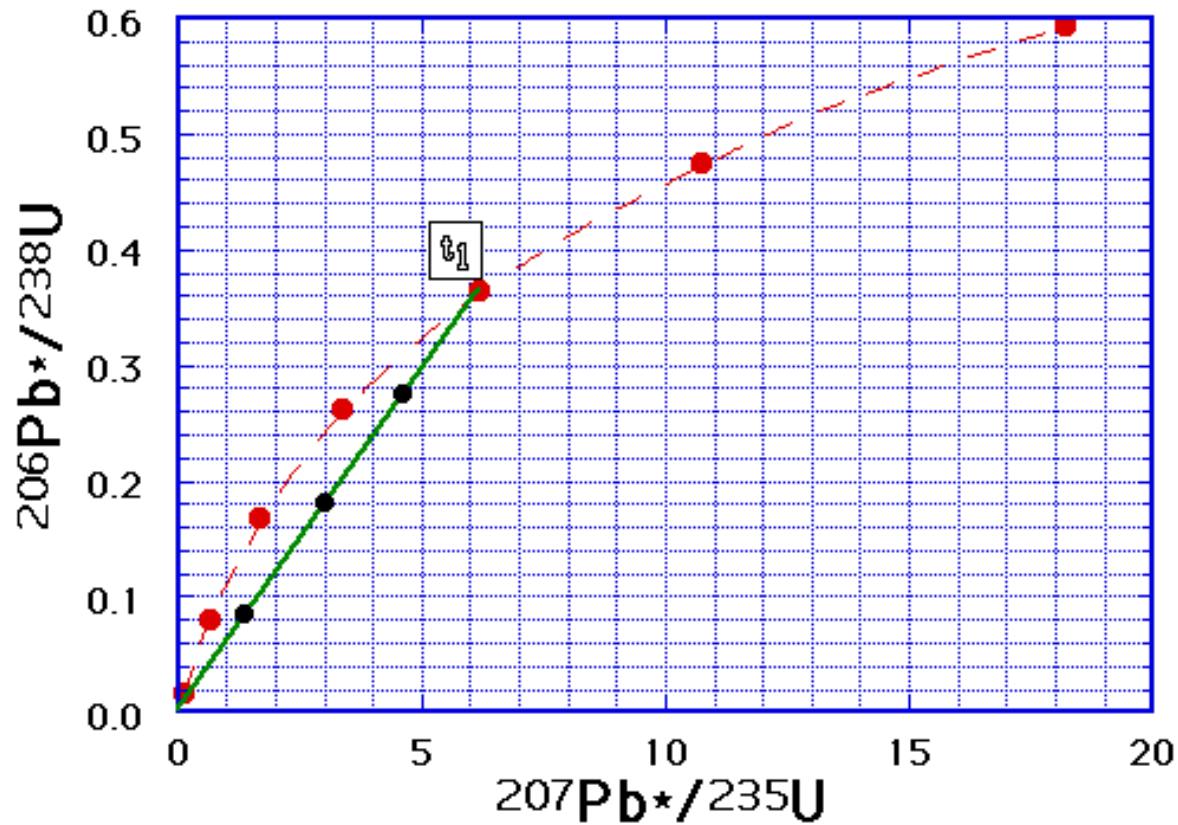
# U-Pb Concordia Diagram

$$^{206}\text{Pb}^*/^{238}\text{U} = (e^{\lambda t_1} - 1), \quad ^{207}\text{Pb}^*/^{235}\text{U} = (e^{\lambda t_2} - 1)$$

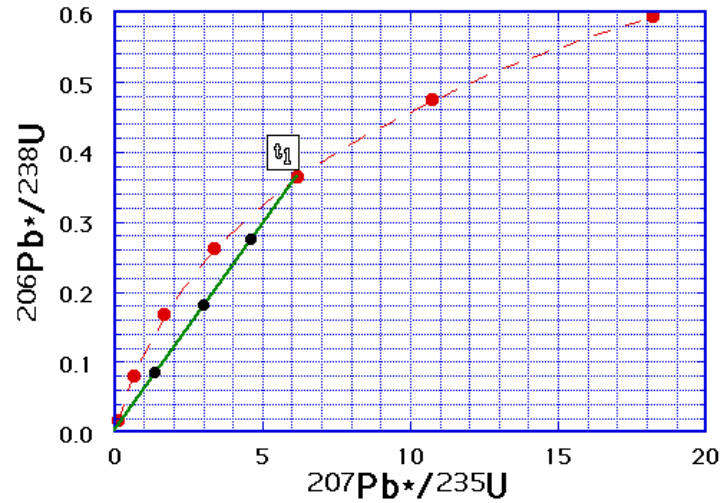


# U-Pb Datierung

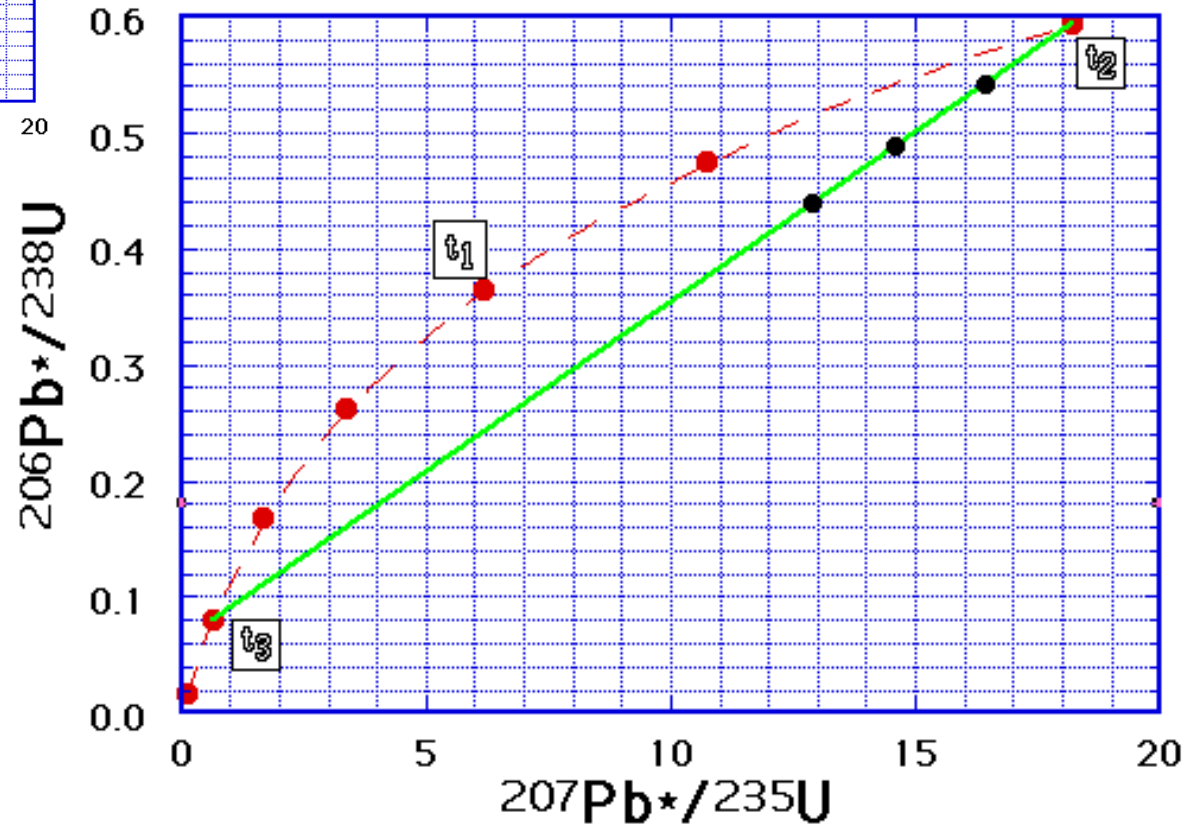
$$^{206}\text{Pb}^*/^{238}\text{U} = (e^{\lambda t} - 1), \quad ^{207}\text{Pb}^*/^{235}\text{U} = (e^{\lambda t} - 1)$$



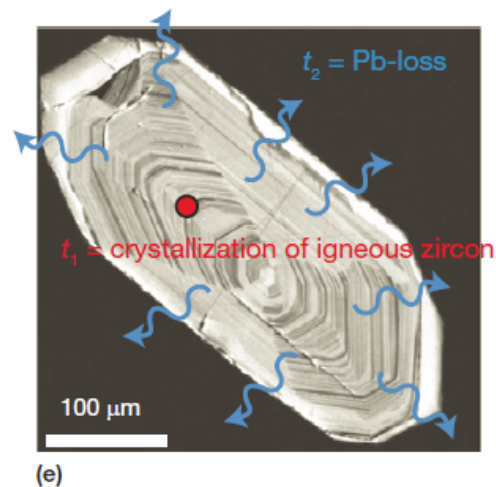
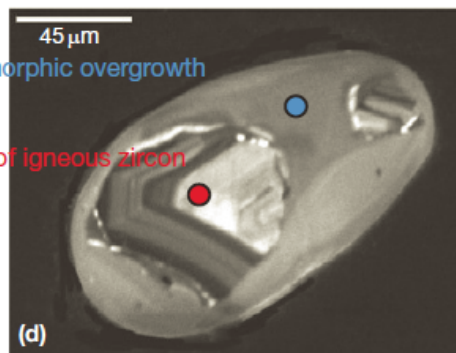
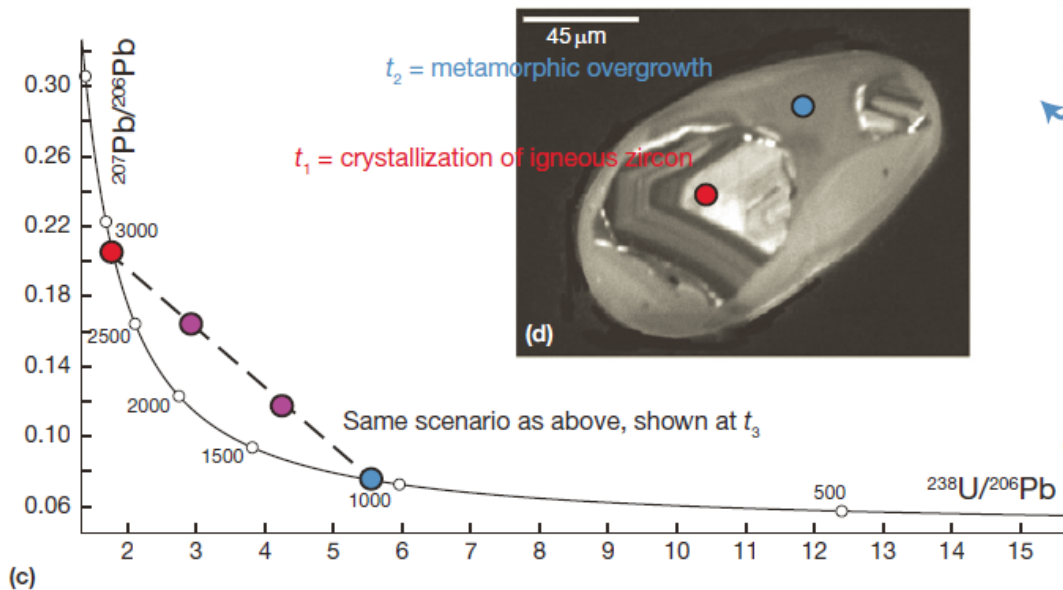
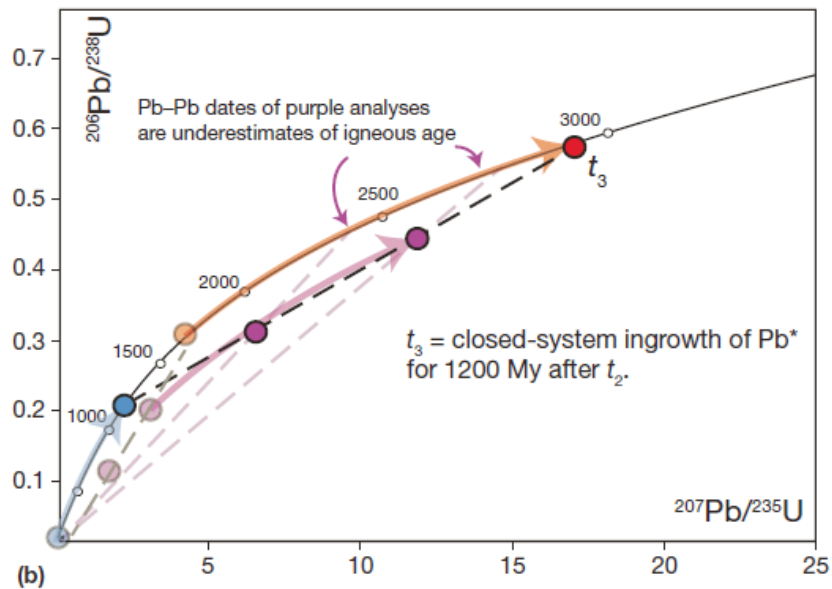
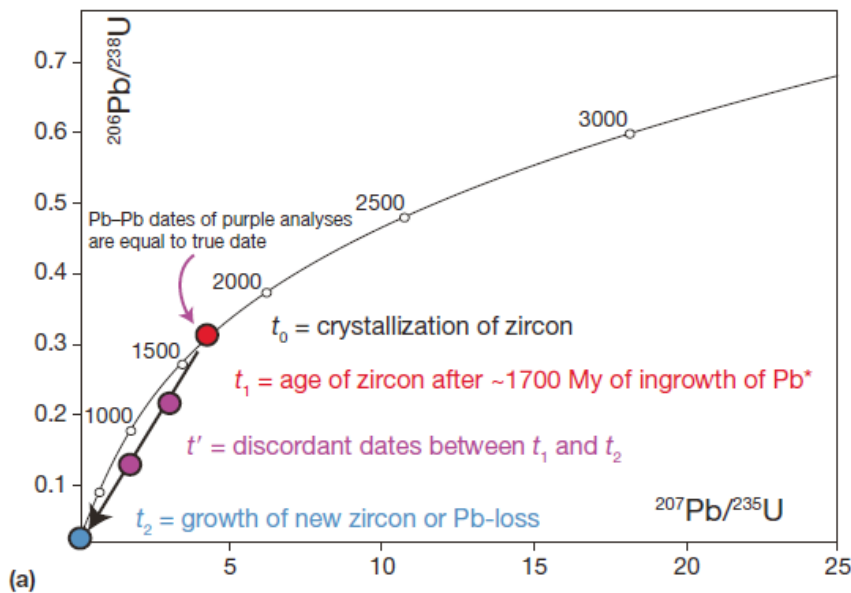
# U-Pb Datierung



$$^{206}\text{Pb}^*/^{238}\text{U} = (e^{\lambda t_1} - 1),$$
$$^{207}\text{Pb}^*/^{235}\text{U} = (e^{\lambda t_2} - 1)$$



# Wetherill & Tera-Wasserburg concordia

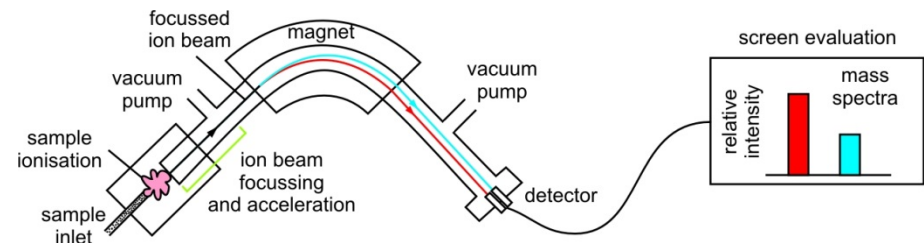


# U-Pb Datierung

## Reinraumlabor



## Massenspektrometer



# Stabile Isotope

**Table 1.2** Characteristic physical properties of  $\text{H}_2^{16}\text{O}$ ,  $\text{D}_2^{16}\text{O}$ , and  $\text{H}_2^{18}\text{O}$

| Property                                | $\text{H}_2^{16}\text{O}$ | $\text{D}_2^{16}\text{O}$ | $\text{H}_2^{18}\text{O}$ |
|---|---------------------------|---------------------------|---------------------------|
| Density (20 °C, in $\text{g cm}^{-3}$ ) | 0.997                     | 1.1051                    | 1.1106                    |
| Temperature of greatest density (°C)    | 3.98                      | 11.24                     | 4.30                      |
| Melting point (760 Torr, in °C)         | 0.00                      | 3.81                      | 0.28                      |
| Boiling point (760 Torr, in °C)         | 100.00                    | 101.42                    | 100.14                    |
| Vapor pressure (at 100 °C, in Torr)     | 760.00                    | 721.60                    |                           |
| Viscosity (at 20 °C, in centipoise)     | 1.002                     | 1.247                     | 1 .056                    |

“isotope effects”:

Differences in chemical and physical properties arising from variations in atomic mass of an element or molecule

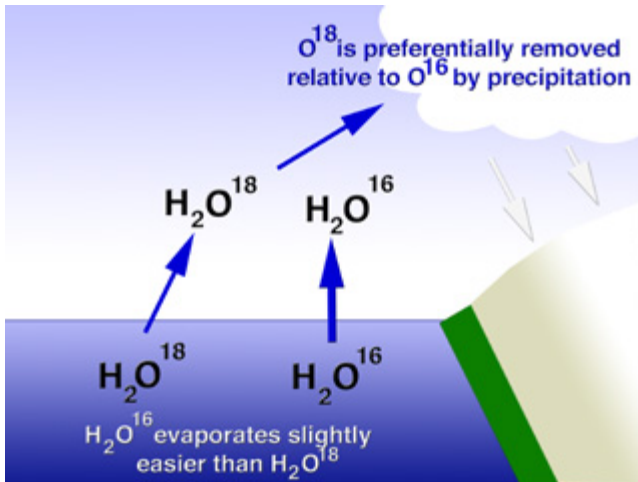
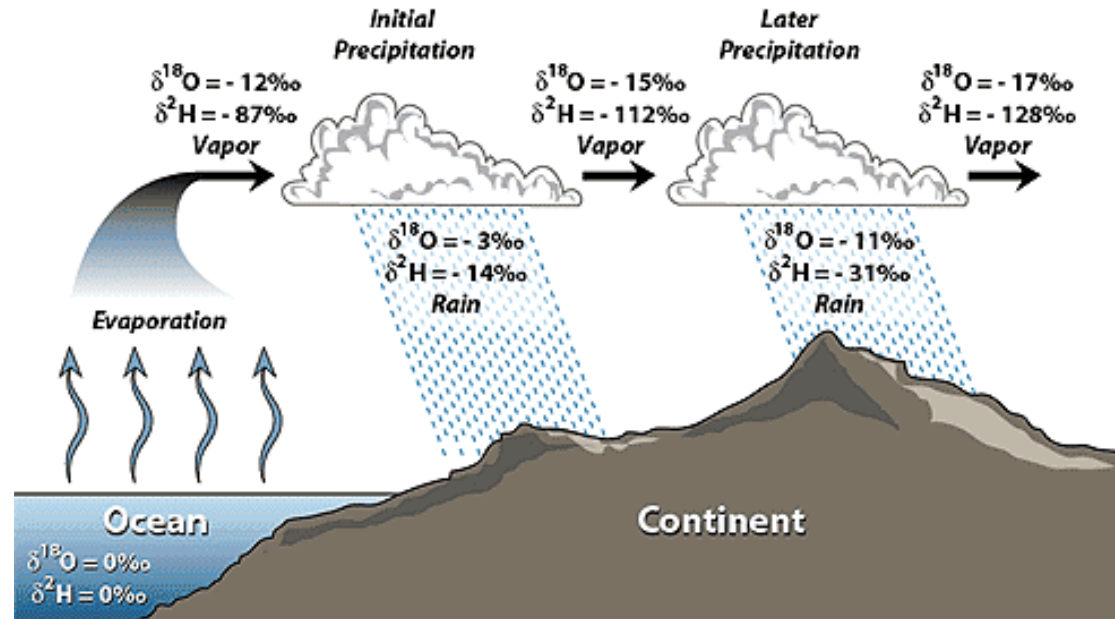


# Stabile Isotope

## Evaporation/precipitation

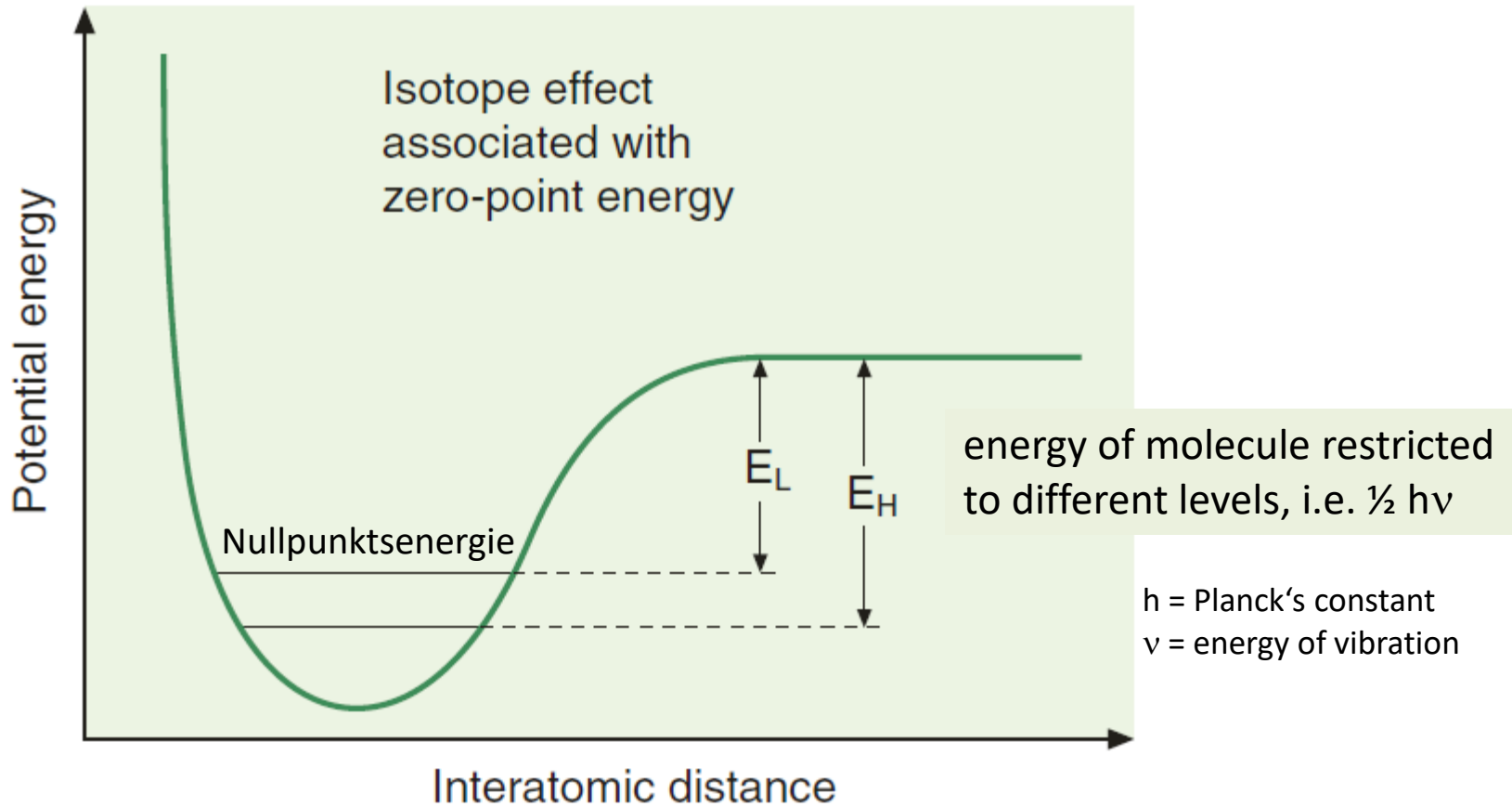
The lighter isotopes evaporate more easily

Heavier isotopes enriched in remaining liquid phase



| Property                              | H <sub>2</sub> O | D <sub>2</sub> O | H <sub>2</sub> <sup>18</sup> O |
|---------------------------------------|------------------|------------------|--------------------------------|
| Density at 20°C [g cm <sup>-3</sup> ] | 0.9982           | 1.1051           | 1.1106                         |
| Melting point [°C]                    | 0                | 3.81             | 0.28                           |
| Boiling point [°C]                    | 100              | 101.42           | 100.14                         |

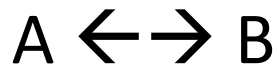
# Stabile Isotope



# Massenabhängige Isotopenfraktionierung

Isotope fractionation during chemical, physical and biological processes:

reversible chemical reaction at equilibrium state



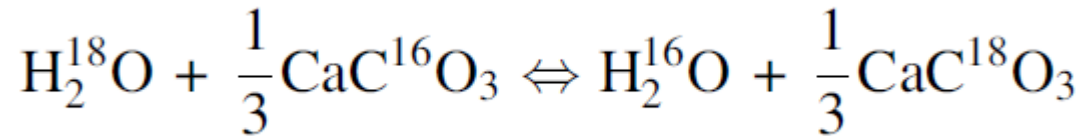
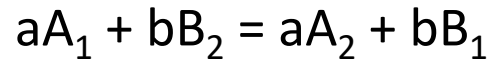
physical changes, phase transitions (e.g. water – vapor), diffusion

biological and biochemical reactions



# Massenabhängige Isotopenfraktionierung

## 1. isotope exchange reactions – example



$$\alpha_{\text{CaCO}_3-\text{H}_2\text{O}} = \frac{\left(\frac{^{18}\text{O}}{^{16}\text{O}}\right)_{\text{CaCO}_3}}{\left(\frac{^{18}\text{O}}{^{16}\text{O}}\right)_{\text{H}_2\text{O}}} = 1.031 \text{ at } 25^\circ\text{C}$$

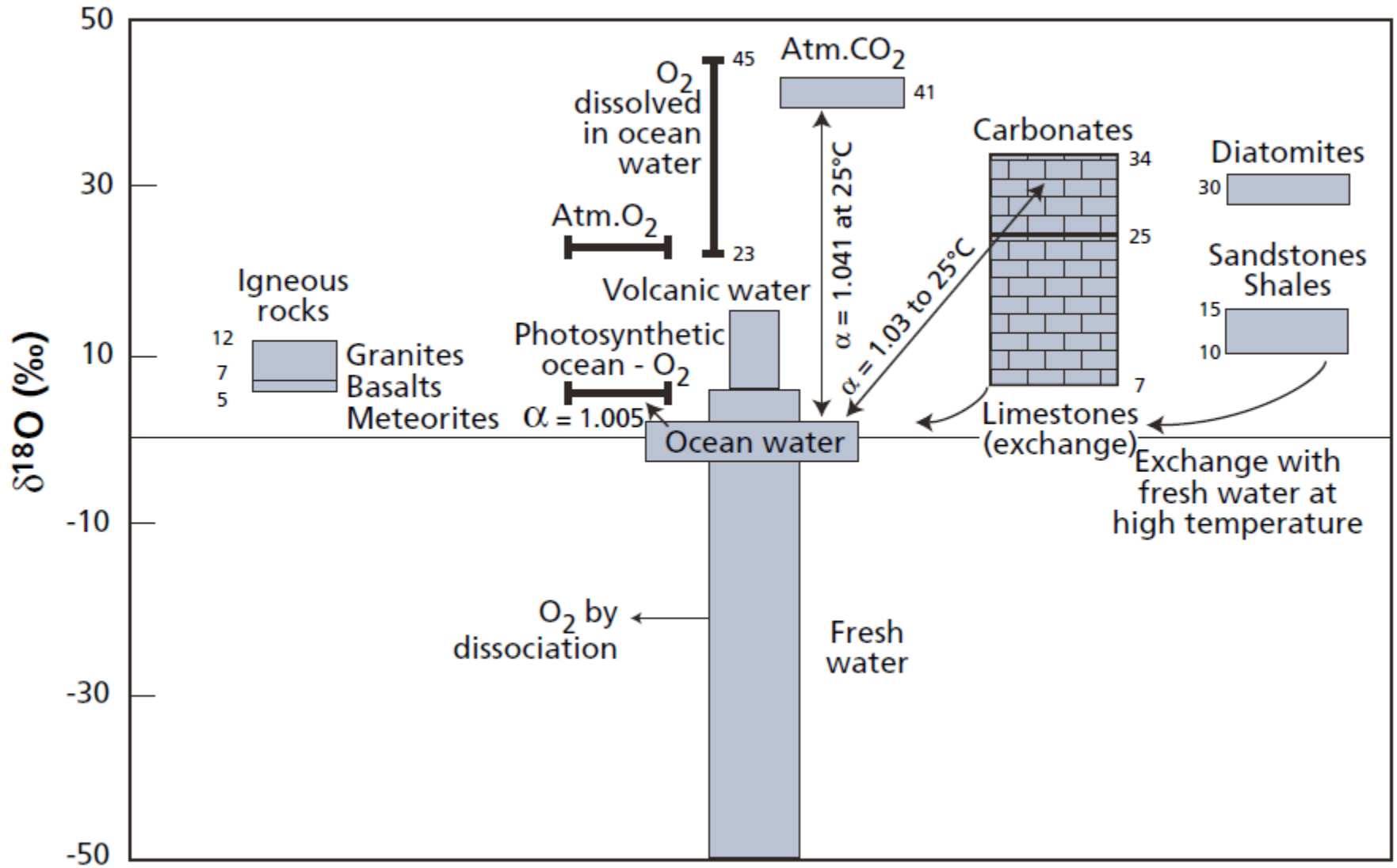
# Massenabhängige Isotopenfraktionierung

## 1. isotope exchange reactions - Delta-value ( $\delta$ )

$$\delta = \left( \frac{\text{sample isotope ratio} - \text{standard isotope ratio}}{\text{standard isotope ratio}} \right) \times 10^3$$

$\delta$  is a relative deviation from a standard, expressed as the number of parts per mil. Isotope ratios are expressed with the heavier isotope in the numerator, e.g.:  $^{18}\text{O}/^{16}\text{O}$ , D/H,  $^{13}\text{C}/^{12}\text{C}$

# Sauerstoff-Isotopenvariationen



# Sauerstoff-Isotopenvariationen

